

1 **Ultraclean layers and optically thin clouds in the stratocumulus to cumulus**

2 **transition: part I. Observations**

3 Robert Wood\*, Kuan-Ting O, Christopher, S. Bretherton, Johannes Mohrmann

4 *Department of Atmospheric Sciences, University of Washington, Seattle, USA*

5 Bruce. A. Albrecht, Paquita Zuidema

6 *Rosenthal School of Marine and Atmospheric Sciences, University of Miami, USA*

7 Virendra Ghate, Chris Schwartz

8 *Argonne National Laboratory, Lemont, USA*

9 Ed Eloranta

10 *University of Wisconsin, Madison, USA*

11 Susanne Glienke, Raymond Shaw

12 *Michigan Technological University, Houghton, USA*

13 Jacob Fugal

14 *Max Planck Institute for Chemistry, Mainz, USA*

15 Patrick Minnis

16 *NASA Langley Research Center, Virginia, USA*

<sup>17</sup> \**Corresponding author address:* Robert Wood, 718 ATG Building, Department of Atmospheric  
<sup>18</sup> Sciences, University of Washington, Seattle, WA 98112, United States of America  
<sup>19</sup> E-mail: robwood2@uw.edu

## ABSTRACT

20 A common feature of the stratocumulus to cumulus transition (SCT) is the  
21 presence of layers in which the concentration of particles larger than  $0.1 \mu\text{m}$   
22 is below  $10 \text{ cm}^{-3}$ . These *ultra clean layers* (UCLs) are explored using air-  
23 craft observations from 14 flights of the NSF/NCAR G-V aircraft between  
24 California and Hawaii. UCLs are commonly located in the upper part of de-  
25 coupled boundary layers, with coverage increasing from only a few percent  
26 within 500 km of the California coast to  $\sim 30\text{-}60\%$  west of  $130^\circ\text{W}$ . Most  
27 clouds in UCLs are thin, horizontally extensive layers containing drops with  
28 median volume radii ranging from  $15\text{-}30 \mu\text{m}$ . Many UCL clouds do not fully  
29 attenuate the G-V lidar, and yet are frequently detected with a 94 GHz radar  
30 with a sensitivity of around  $-30 \text{ dBZ}$ . Satellite data indicate that UCL clouds  
31 have visible reflectances of  $\sim 0.1\text{-}0.2$  and are often quasi-laminar, giving them  
32 a veil-like appearance. These veil clouds exist for 1-3 hours or more, are asso-  
33 ciated with mesoscale cumulus clusters, and likely grow by spreading under  
34 strong inversions. Active updrafts in the Cu clouds have droplet concentra-  
35 tions of  $\sim 25\text{-}50 \text{ cm}^{-3}$ . Collision-coalescence in the Cu, and later sedimenta-  
36 tion in the thinner UCL clouds, are likely the key processes removing droplets  
37 in UCL clouds. UCLs are relatively quiescent, and a lack of mixing with dry  
38 air above and below the cloud may help to explain their longevity. The very  
39 low and highly variable droplet concentrations in UCL clouds, together with  
40 their low geometrical thickness, makes these clouds particularly challenging  
41 to represent in large scale models.

## 42 **1. Introduction**

43 Assessment of the physical factors controlling the coverage and albedo of marine boundary  
44 layer (MBL) clouds remains a pressing challenge. The speed at which the stratocumulus (Sc) to  
45 Cumulus (Cu) transition (SCT) occurs in air masses downstream of the eastern subtropical ocean  
46 basins determines the albedo of the Tropics, and similar cloud transitions in post-frontal air masses  
47 are important for determining midlatitude storm track albedo. These transitions in cloudiness are  
48 poorly represented in climate models, which typically show stratocumulus break-up occurring too  
49 quickly in both the subtropics (Teixeira et al., 2011) and the midlatitudes (Williams et al., 2012;  
50 Bodas-Salcedo et al., 2014). In addition to producing too few low clouds, model clouds are too  
51 reflective, a pathology commonly referred to as the “too-few, too-bright” problem (Nam et al.,  
52 2012; Engstrm et al., 2015).

53 In general, the SCT initiates in response to MBL deepening and decoupling as surface fluxes  
54 increase over warmer waters (Bretherton and Wyant, 1997; Wyant et al., 1997). As several studies  
55 have shown (Martin et al., 1995; Zhou et al., 2015), stratocumulus cloud breakup is not an imme-  
56 diate response to MBL decoupling, but can be delayed by as much as 1-3 days, with cloud cover  
57 often remaining above 50% until 500-2000 km downsteam of decoupling onset (Zhou et al., 2015).  
58 Given the importance of the SCT, it is remarkable that few dedicated aircraft observations have  
59 sampled the MBL in the breakup region owing to the typically remote location of this phenomenon  
60 and the limited range of most research aircraft. Early transition experiments with shipborne obser-  
61 vations (e.g., the 1969 Airmass Transition Experiment, ATEX, Augstein et al., 1973; Stevens et al.,  
62 2001) provided important characterization of the vertical thermodynamic structure of the MBL un-  
63 der transition conditions, but little if any information on cloud structure and dynamics. Aircraft  
64 observations made during the Atlantic Stratocumulus Transition Experiment (ASTEX) in 1992

65 provide one of the most widely used datasets for understanding the SCT and are perhaps the first  
66 measurements of cloud microphysical properties in the SCT. The two Lagrangian transition exper-  
67 iments carried out in ASTEX showed that drizzle occurred with sufficient frequency and rate to be  
68 important in MBL energy and moisture budgets (Bretherton and Pincus, 1995; Bretherton et al.,  
69 1995). Surface-based radar measurements in ASTEX also highlighted the frequent occurrence of  
70 precipitating cumulus clusters with horizontal scales of  $\sim 10$  km rising into and interacting with  
71 stratocumulus clouds at the top of the decoupled MBL (Miller and Albrecht, 1995). Martin et al.  
72 (1995) showed that the stratocumulus deck is locally thickened in regions of interaction with ris-  
73 ing Cu, consistent with the Cu providing a source of moisture to the decoupled cloud layer aloft.  
74 However, the Cu also drive entrainment of dry air from the free troposphere, which eventually  
75 serves to dry the upper MBL and dissipate the Sc (Krueger et al., 1995; Wyant et al., 1997). The  
76 timescale over which drying happens, i.e., the time between decoupling and stratocumulus break  
77 up, likely depends upon the relative amount of free tropospheric entrainment that is achieved for  
78 a given Cu mass flux, the inversion strength (which influences both entrainment and the spreading  
79 of Cu below the inversion), the free-tropospheric humidity, and the rate at which the upper MBL  
80 loses condensate through precipitation and droplet sedimentation. The processes causing drying  
81 therefore are expected to depend upon the nature of the clouds and turbulent dynamics in the upper  
82 MBL.

83 Large eddy simulations (LES) indicate that the SCT is sensitive to cloud microphysical pro-  
84 cesses in addition to large scale meteorological controls. Sandu and Stevens (2011) found that  
85 reducing the cloud droplet concentration  $N_d$  (assumed to be constant throughout the domain and  
86 over the simulation time) from 100 to  $33 \text{ cm}^{-3}$  increases precipitation efficiency and slows down  
87 the growth of the MBL, but hastens the rate of cloudiness reduction. Although not specifically  
88 transition experiments, previous short-duration LES studies have shown a strong dependence of

89 stratocumulus cloud cover on  $N_d$  (Ackerman, 2003), especially when  $N_d$  falls below  $50 \text{ cm}^{-3}$ .  
90 This result is broadly consistent with the Sandu and Stevens (2011) study, pointing to the need to  
91 understand what controls  $N_d$  in the SCT.

92 Several cases of spatial transitions from closed to open cellular convection over the southeastern  
93 Pacific Ocean were observed during the VOCALS Regional Experiment (Terai et al., 2014), and  
94 showed that within regions of open cells, active Cu clouds that draw aerosol from the surface mixed  
95 layer have  $N_d$  values approximately three times larger than the layer of thin stratus/stratocumulus  
96 at the top of the MBL into which the Cu are detraining. Further,  $N_d$  in the stratus layer of the open  
97 cell region was found to be extremely low ( $5 \text{ cm}^{-3}$  in the mean, over five cases), prompting its  
98 characterization as an *ultra clean layer* (UCL), a term first used for clean, thin unpolluted layers  
99 between aerosol-rich polluted layers over southern Africa (Hobbs, 2003). In one open cell case  
100 during VOCALS,  $N_d$  in the most active Cu updrafts was over an order of magnitude greater than  
101 that in the relatively quiescent detrained stratus layer (Wood et al., 2011). Other observational  
102 studies have found cases with extremely low  $N_d$  and cloud condensation nuclei concentrations in  
103 the subtropical MBL (Stevens et al., 2005; Sharon et al., 2006; Petters et al., 2006; Wood et al.,  
104 2008), in midlatitude marine cold air outbreaks (Field et al., 2014) and in the Arctic (Mauritsen  
105 et al., 2011). Combining data from satellite with ship cruises between California and Hawaii,  
106 Painemal et al. (2015) showed using that the correlation of satellite  $N_d$  and surface-based accu-  
107 mulation aerosol concentration was 0.9 east of  $145^\circ\text{W}$ , but dropped to 0.5 in the SCT west of  
108 this, consistent with a decoupling of cloud microphysics from near-surface aerosol and surface  
109 aerosol sources in the SCT. This would be consistent with an aerosol sink in the upper PBL that is  
110 uncorrelated with near-surface aerosol.

111 The dramatic variability of  $N_d$  even within the same cloud system raises the question of whether  
112 fixing  $N_d$  (or aerosol concentration) across an entire model domain is appropriate for the study of

113 the SCT. Recent LES studies (e.g., Kazil et al., 2011; Berner et al., 2013) have allowed  $N_d$  and  
114 aerosols to vary spatiotemporally in response to coalescence scavenging and Brownian scavenging,  
115 surface particle fluxes, entrainment, and advection. These studies have been able to reproduce with  
116 some fidelity the vertical and horizontal variability of  $N_d$  and aerosol concentration in the MBL  
117 seen in the open cell cases observed in VOCALS. A dramatic reduction in  $N_d$  due primarily by  
118 coalescence scavenging appears to be a universal feature of the transition from closed to open  
119 cells, although the level of sophistication in the aerosol budget necessary to faithfully reproduce  
120 the cloudiness transition is unclear. What is also unclear is the extent to which closed to open  
121 cell transitions and specifically the formation of UCLs are representative of the behavior of SCTs  
122 observed elsewhere.

123 This study explores the nature of clouds and aerosols in the upper MBL occurring in the SCT,  
124 with specific focus on understanding the relationship between ultra-clean layers (UCLs) and  
125 frequently-occurring optically thin clouds in these layers. The primary tool is aircraft observa-  
126 tions from fourteen flights of the NSF/NCAR G-V aircraft during the Cloud System Evolution in  
127 the Trades (CSET) field program. CSET sampled the MBL and lower free troposphere extend-  
128 ing across the entire region from the Californian coast ( $\sim 38^\circ\text{N}$ ,  $123^\circ\text{W}$ ) to Kona, Hawaii ( $19^\circ\text{N}$ ,  
129  $156^\circ\text{W}$ ), a distance of 3700 km. Trajectory calculations indicate that air-masses take approxi-  
130 mately five days to advect this distance (Sandu et al., 2010), offering unprecedented observations  
131 of the complete transition from shallow marine stratocumulus and stratus, to deep well-mixed and  
132 then decoupled Sc, through to a largely trade Cu environment at Hawaii, albeit one with 40%  
133 coverage by optically thin upper PBL stratiform clouds. Section 2 describes the observational  
134 datasets used and the methodological approach. Section 3 examines the statistical occurrence of  
135 both clear and cloudy UCLs including their frequency, altitude, and geographical distribution.  
136 Section 4 documents properties of both cloudy and clear UCLs using *in situ* G-V data. Section 5

137 explores several case studies from CSET, combining *in situ* and remote sensing data from the G-V  
138 in conjunction with satellite remote sensing, to provide further insights into the microphysical and  
139 macrophysical structure of UCLs and optically thin clouds. Section 6 documents the statistical  
140 properties of optically thin MBL clouds observed in CSET, relating their vertical structure to that  
141 of cloudy UCLs. Finally, Section 7 introduces hypotheses for how UCLs are formed and discusses  
142 some of the implications of the results presented in this study.

## 143 **2. Data and methods**

### 144 *a. Instruments and sampling*

145 Data from the Cloud System Evolution in the Trades (CSET) field campaign (Albrecht, 2017)  
146 constitute the primary observational dataset used in this study. The NSF/NCAR Gulfstream-V  
147 High-performance Instrumented Airborne Platform for Environmental Research (G-V HIAPER)  
148 aircraft sampled marine air masses between Sacramento, California (38.6°N, 121.5°W) and Kona,  
149 Hawaii (19.6°N, 156.0°W). The G-V was instrumented with numerous instruments, which are  
150 detailed in Albrecht (2017) and are only described briefly here. In addition to measurements  
151 of the atmospheric state variables (temperature, pressure, water vapor, and winds including high  
152 frequency turbulent components at 20 Hz), other instruments used in this study include:

- 153 • an Ultra High Sensitivity Aerosol Spectrometer (UHSAS) that measures aerosol particles  
154 with diameters from 60-1000 nm, with 100 size bins. Here, we use only the most reliable size  
155 bins from 100-1000 nm and define  $N_d$  as the concentration of particles larger than 100 nm.  
156 UHSAS data are used only in cloud-free conditions;
- 157 • Cloud droplet probe, measuring the cloud droplet size distribution from 1-25  $\mu\text{m}$  radius (with  
158 a 1  $\mu\text{m}$  bin width), and a fast 2D-C optical array probe to provide a size distribution for larger

159 drops between 37.5 and 800  $\mu\text{m}$  radius with a bin width of 12.5  $\mu\text{m}$  radius. The concentration  
160 of drops measured by the CDP is termed the droplet concentration  $N_d$ . The CDP and 2D-C  
161 data are combined to produce a single cloud-drizzle size distribution from 1-800  $\mu\text{m}$ , with  
162 spline interpolation in log-log space used to fill the gap region between the upper size bin of  
163 the CDP and the smallest useable bin of the 2D-C; in addition to the CDP and 2D-C probes,  
164 the G-V also flew a holographic cloud droplet detector (HOLODEC) (Fugal and Shaw, 2009)  
165 that uses holographic reconstruction to measure droplets spanning the poorly measured gap  
166 region.

- 167 ● King hotwire probe to measure liquid water content;
- 168 ● A VCSEL water vapor hygrometer and a Rosemount temperature sensor;
- 169 ● HIAPER Cloud Radar (HCR), a 94 GHz Doppler radar that can sample either in the nadir  
170 or zenith directions (Vivekanandan et al., 2015). Nadir sampling is used when the aircraft  
171 is above cloud or in cloud layers, and zenith sampling is used when the aircraft is below the  
172 cloud level;
- 173 ● High Spectral Resolution Lidar (HSRL), a 532 nm lidar (Grund and Eloranta, 1991) that can  
174 sample either in near-nadir or near-zenith directions. The lidar is manually turned from nadir  
175 to zenith sampling and the pointing direction is set to be almost the same as that for the HCR,  
176 except that the HSRL is designed to be pointed  $4^\circ$  laterally off-nadir/zenith to avoid specular  
177 reflections from ice crystals.

178 We use all available data from fourteen of the research flights (RF02 through RF15) that were  
179 designed to sample airmasses transitioning from Sc to Cu. The outbound flights (Sacramento  
180 to Kona) comprised a transit leg (typically at 6-8 km altitude) from Sacramento to an offshore

181 location containing Sc clouds relatively close to the Californian coast. The G-V then profiled  
182 down to conduct low-level sampling of the PBL and lower free troposphere (FT) for a horizontal  
183 distance of typically 1500-2000 km before ascending again to  $\sim 6$  km for the transit to Kona.  
184 Low-level sampling constituted approximately half of the flight. The return flight from Kona to  
185 Sacramento was conducted two days after the outbound flight and was designed to sample the  
186 locations of the Lagrangian-advected PBL airmasses that had moved toward Hawaii over the two  
187 days between the outbound and return flights. The return flight consisted of approximately the  
188 same amount of low level sampling as was carried out on the outbound flight. The locations of the  
189 flights are shown in (Albrecht, 2017). The low-level sampling strategy consisted of repeated flight  
190 modules comprising a 10 min near-surface straight and level leg ( $\sim 140$  m above the surface), a  
191 10 min straight and level leg 100 m above the cloud base (either the Sc cloud base for overcast  
192 conditions or the Cu cloud base for broken cloud conditions), a sawtooth leg porpoising from  
193 approximately 100 m below cloud top to 400 m above cloud top (with three upward inversion  
194 crossings and two downward), a 10 min straight and level leg in the FT approximately 1000 m  
195 above cloud top, and then a profile from this level down to the near-surface (140 m altitude). This  
196 sampling strategy was designed to provide somewhat representative horizontal and vertical *in situ*  
197 sampling of the PBL and the lower FT, as well as good opportunities for remote sensing of clouds  
198 and precipitation.

### 199 *b. UCL Definition*

200 This study focuses upon the nature and occurrence of ultra clean layers (UCLs) over the subtrop-  
201 ical northeastern Pacific Ocean. We use G-V *in situ* cloud and aerosol properties to identify UCLs  
202 and to characterize their physical properties. Over the southeastern Pacific (Terai et al., 2014),  
203 UCLs were found in all cases of mesoscale open cellular convection, and in these cases typically

204 assumed the form of horizontally extensive layers in the upper PBL below the trade inversion.  
205 These layers typically consist of patches of clear air and patches of cloudy air. In cloud-free con-  
206 ditions, UCLs were marked by very low concentrations of accumulation mode aerosol ( $N_a$  defined  
207 in Terai et al. (2014) and here as particles with diameters larger than  $0.1 \mu\text{m}$ ), whereas under  
208 cloudy conditions UCLs can be identified by very low cloud droplet concentrations ( $N_d$ , droplets  
209 with radii larger than  $1 \mu\text{m}$ ). Defining UCLs requires the choice of thresholds, and this is some-  
210 what arbitrary. We note that mean concentrations ( $N_a$  or  $N_d$  in the upper PBL during cases of open  
211 cells over the southeastern Pacific were generally below  $10 \text{ cm}^{-3}$  (Terai et al., 2014), which is  
212 5-25 times lower than typical mean concentrations found in previous field studies of marine stra-  
213 tocumulus (see e.g., Martin et al., 1994; Miles et al., 2000). There have been very few studies that  
214 have noted such low particle concentrations, yet as we shall see these occur relatively frequently  
215 in the SCT. We therefore set a threshold for  $N_a$  or  $N_d$  of  $10 \text{ cm}^{-3}$  to distinguish UCLs from “non  
216 UCL” samples.

217 The procedure for classifying samples is as follows. We begin with G-V aircraft data sampled  
218 at 1 Hz, or roughly 150 m along the flight direction. This is not necessarily the native frequency  
219 of the measurements, but placing all the measurements for each flight into a single universal time  
220 frame allows the use of multiple instruments and criteria to classify air samples. We first classify  
221 each 1 Hz sample as into one of the following: (a) clear samples, (b) cloudy samples, or (c)  
222 samples containing precipitation under subsaturated conditions.

- 223 • (a) Clear samples must satisfy the following: liquid water content measured using the CDP  
224  $q_{l,CDP}$  must be lower than  $0.01 \text{ g kg}^{-1}$ , liquid water content  $q_{l,2DC}$  measured using the 2D-C  
225 probe must be lower than  $0.001 \text{ g kg}^{-1}$ , and the total drizzle drop concentration  $N_{2DC}$  mea-  
226 sured with the 2D-C probe must not exceed  $1 \text{ liter}^{-1}$ ;

227 • (b) Cloudy samples must have  $q_{l,CDP} > 0.01 \text{ g kg}^{-1}$  and must have relative humidity exceed-  
228 ing 95%.

229 • (c) Samples containing precipitation under unsaturated conditions have either  $q_{l,CDP} >$   
230  $0.01 \text{ g kg}^{-1}$  or  $q_{l,2DC} > 0.001 \text{ g kg}^{-1}$ , or  $N_{2DC} > 1 \text{ liter}^{-1}$ , and must have relative humidity  
231 below 95%.

232 Following classification of samples into cloud, clear and unsaturated precipitation, each of the  
233 cloudy and clear 1 Hz samples is then classified as being either a UCL sample or a non-UCL  
234 sample using the following criterion

235 • (a) To be classified as a clear UCL, the 1 Hz sample must be classified as clear and further  
236 must have  $N_a < 10 \text{ cm}^{-3}$ ;

237 • (b) To be classified as a cloudy UCL, the 1 Hz sample must be classified as cloud-containing  
238 and must have  $N_d < 10 \text{ cm}^{-3}$ .

239 Once each 1 Hz sample has been classified, statistical analysis is performed as described in the  
240 next section.

### 241 3. UCL occurrence statistics

242 Longitude-height cross sections of UHSAS  $N_a$  from RF02 to RF15 (Fig. 1) show that very low  
243 concentrations ( $< 10 \text{ cm}^{-3}$ ) commonly occur, especially west of  $\sim 135^\circ\text{W}$ . It is rare that these  
244 layers of very low  $N_a$  extend to the surface but instead tend to be concentrated at altitudes of 1-  
245 2.5 km, with the deeper cases mostly occurring in the western half of the region. Most flights  
246 indicate the presence of UCLs ( $N_a < 10 \text{ cm}^{-3}$ ) extending some distance between California and  
247 Hawaii, but there is considerable variability between cases. For example, RF07 shows that UCLs  
248 were sampled on every leg over the  $\sim 2000 \text{ km}$  distance from  $132\text{-}152^\circ\text{W}$ , and were growing in

249 height westward with the growth in the PBL. Interestingly, RF02 and RF03, which show few  
250 UCLs, were severely impacted by biomass burning aerosol from forest fires in British Columbia  
251 that entrained into the PBL.

252 Using the classification described in the previous section, we analyze clear and cloudy UCLs  
253 separately. As Fig. 1 shows, the low level sampling in CSET extended from 128°W, some distance  
254 upstream of the stratocumulus maximum (climatologically at ~135°W), to 152°W, close to Hawaii  
255 and well into the SCT breakup region. Mean low cloud cover over this longitudinal range falls  
256 from 0.7 to 0.4 (see Section 6). Our aim in this study is to examine how UCL occurrence frequency  
257 varies both with height and with longitude. To do this, we aggregate 1 Hz flight samples flight-by-  
258 flight into 5° longitude and 150 m altitude bins. The longitude bins are chosen to be sufficiently  
259 wide that they encompass an entire flight module (see section 2), providing good vertical statistics  
260 in the PBL and lower FT within a single longitude bin. Separately for clear and cloudy samples,  
261 we determine the fraction of samples in each flight/longitude/height bin that are classified as UCL  
262 samples.

263 An aircraft flying along a path without backtracking can only provide *in situ* measurements at a  
264 single height for any given longitude. In general, this precludes the measurement of vertically in-  
265 tegrated quantities such as cloud cover or UCL cover using *in situ* aircraft data. For example, the  
266 aircraft may happen to be sampling a clear region above clouds when there are clouds below. How-  
267 ever, the rigid flight sampling strategy used in CSET, most importantly the use of predetermined  
268 flight modules along a predetermined flight track, with leg height choices objectively chosen to  
269 ensure good sampling of the PBL and lower FT, means that each flight module can provide useful  
270 information on the fraction of samples that are within UCLs. We argue that whereas the problem  
271 of estimating cloud cover with *in situ* G-V data is not well posed, an assessment of the *fraction* of  
272 clear sky and clouds that are UCLs is less biased and is statistically meaningful.

273 For a first look at vertical UCL statistics, we average the UCL fraction profiles over the CSET  
274 flights and longitude bins with sufficient samples to give vertical profiles of the fraction of clear  
275 and cloudy samples respectively that are diagnosed to be UCLs. To be considered sufficient,  
276 there need to be at least 20 samples in a given flight/longitude/height bin to be considered in the  
277 average. The frequency of both cloudy and clear UCLs increases with height from 0.5 km to 1.5-  
278 2 km altitude (Fig. 2). The highest fraction of clear samples classified as UCLs is approximately  
279 20%, occurring at 1.5-1.8 km altitude. Cloudy samples are almost twice as likely to be classified as  
280 UCLs compared with clear samples, with values of 30-50% at 2 km but the vertical trends for each  
281 are quite similar. Above 2 km the clear UCL fraction falls steadily to  $\sim 10\%$  at 3 km altitude. The  
282 sampling at altitudes above 3km was too sparse to produce meaningful statistics. Below 0.5 km,  
283 there were few clouds sampled in CSET, and only 1-2% of clear samples were found to be UCLs.  
284 Above 2.0 km, there were an insufficient number of flights sampling clouds to produce statistically  
285 representative cloudy UCL fractions.

286 We next estimate the geographical variation of UCLs between California and Hawaii using the  
287 following approach. We seek the conditional probability (given a cloud or clear column) that the  
288 column contains a UCL *at some level*, assuming maximum vertical overlap. For each flight and  
289  $5^\circ$  longitude bin with a sufficient number of samples, we identify the height bin containing the  
290 largest UCL fraction (this is done separately for cloudy and clear UCLs). Then, we average these  
291 maximum fractions over all the flights to give an estimate of the probability that a given longitude  
292 bin contains a UCL somewhere in the column. We accept that the aircraft *in situ* sampling is  
293 sparse and that this estimate of mean column UCL fraction is likely to be biased, but it is not  
294 obvious whether this will lead to a high or low bias. Assuming random overlap or maximum-  
295 random overlap rather than maximum overlap (see e.g., Morcrette and Fouquart, 1986) will lead  
296 to larger estimates of the column fraction, but we note that for many of the CSET cases, there

297 is a single layer (or two adjacent layers) with UCL fractions substantially higher than any of the  
298 other layers. In these cases, there is little difference between estimates assuming random and  
299 maximum overlap. On the other hand, the rather large ( $5^\circ$  longitude) horizontal distance over  
300 which we aggregate aircraft data coupled with the limited number of samples in some height  
301 bins, may generate UCL column fraction estimates that are too high. For example, for any given  
302 flight/longitude/height bin, we only require twenty 1 Hz samples to constitute sufficient sampling.  
303 With an aircraft climb/descent rate of 500 m per minute, it takes approximately 20 seconds to  
304 profile through 150 m altitude (i.e., the thickness of our height bins). If all 20 seconds of this  
305 sampling consisted of e.g., clear UCL samples, then the clear UCL fraction would be unity and  
306 this would be carried through as the column UCL fraction for the entire longitude bin for that  
307 flight. If we increase the number of samples required before a particular bin is considered, then  
308 we exclude an increasing number of height bins from the averaging and reduce our reliability of  
309 capturing the true vertical profile.

310 Figure 3 shows the CSET estimated conditional probability (given cloudy or clear columns)  
311 of finding a UCL in a given column, plotted as a function of longitude along the CSET sampling  
312 region from California to Hawaii. Here, we refer to this probability as the fractional UCL coverage.  
313 UCL coverage is only few percent east of  $130^\circ\text{W}$  consistent with most prior observations in the  
314 region within about 500 km of the California coast where very low concentrations of aerosol  
315 and/or cloud droplets are relatively rare. A caveat is that only three of the CSET flights in this  
316 region had sufficient sampling to contribute to this average, and so our statistics are rather poor in  
317 this region. And whereas there have been some notable studies of very low aerosol and/or cloud  
318 droplet concentration clouds in the region east of  $130^\circ\text{W}$  (e.g., Hindman et al., 1994; Petters et al.,  
319 2006), a survey of the published literature of studies in the near-coastal Californian stratocumulus  
320 deck (see e.g., Hudson and Xie, 1999; Miles et al., 2000; Lu et al., 2007) show that very low

321 CCN conditions are relatively rare. This is consistent with a strong offshore gradient in cloud  
322 droplet concentration measured as seen from satellites, where near-coastal mean values exceed  
323  $100 \text{ cm}^{-3}$  (Wood et al., 2012). West of  $130^\circ\text{W}$ , both clear and cloudy the UCL frequencies rise  
324 rapidly (Fig. 3) to values in the range of 40-50% between  $130\text{-}140^\circ\text{W}$ , and reach as high as 60%  
325 by  $150^\circ\text{W}$ . The westward increase for clear UCLs is somewhat less noisy than for cloudy UCLs,  
326 which also occur somewhat less frequently than clear UCLs west of  $140^\circ\text{W}$ . This is because the  
327 overall frequency of clouds decreases westward as the extensive stratocumulus clouds break up  
328 (thus making statistics less reliable), but may also reflect a general tendency for clear UCLs to be  
329 more horizontally extensive than cloudy UCLs. Thus even as the probability of finding a UCL  
330 cloud in any given cloudy sample is higher than it is for the clear counterparts (Fig. 2), clear UCLs  
331 tend to occur somewhere in the column more frequently than cloudy UCLs.

#### 332 **4. In situ properties of UCLs**

##### 333 *a. Aerosol and cloud properties*

334 In this section, we use the G-V *in situ* data to examine vertically-resolved characteristics of  
335 cloudy and clear UCLs. In clear UCLs, concentrations of aerosol particles with diameters exceed-  
336 ing  $0.1 \mu\text{m}$  ( $N_a$ ) decrease with height from values only slightly smaller than the threshold used  
337 to define UCLs ( $10 \text{ cm}^{-3}$ ) in the surface mixed layer below  $\sim 600 \text{ m}$  to values of  $2\text{-}5 \text{ cm}^{-3}$  in  
338 the height range (1-2 km) where most UCLs occur (Fig. 4(a)). Remarkably, mean cloud droplet  
339 concentrations  $N_d$  shows very similar vertical behavior up to 2 km altitude. The strong connection  
340 between the vertical profiles of  $N_a$  and  $N_d$  is consistent with the hypothesis that clear UCLs essen-  
341 tially result from the evaporation of UCL clouds, which being few in number results in a very low  
342 return of cloud-forming aerosol particles to the clear sky. An alternative hypothesis could be that

343 some UCL clouds initially form upon saturation of clear UCLs and the associated activation of a  
344 very low concentration of accumulation mode aerosol. In either case, the question of what drives  
345 the very low concentrations in the first place must be addressed.

346 Although accumulation mode aerosol concentrations are, by definition, very low in UCLs, the  
347 median concentration  $N_{CN}$  of condensation nuclei (particles with diameters larger than 10 nm) in  
348 clear UCLs is 50-90  $\text{cm}^{-3}$  (Fig. 4(b)), i.e., over an order of magnitude higher than  $N_a$ . Further,  
349 median  $N_{CN}$  in clear UCLs is relatively independent of UCL altitude from the surface up to 2 km.  
350 In contrast, median  $N_{CN}$  for all clear samples exceeds 130  $\text{cm}^{-3}$  at all levels, increasing to approx-  
351 imately 200  $\text{cm}^{-3}$  in the lower free troposphere. Clear UCLs have anomalously high mean relative  
352 humidity exceeding 80% at all levels and closer to 90% in the most common UCL height range  
353 of 1-2 km (Fig. 4(b)). Clear UCLs at altitudes above 2 km show higher mean CN concentrations  
354 and lower mean RH than those layers below this, and some of these UCLs are actually in the free  
355 troposphere (FT). Probability distribution functions for clear samples below 3 km altitude (Fig. 5)  
356 show that 60% of UCLs have RH exceeding 85%. In contrast, only 28% of all clear samples  
357 have such high RH. This figure also demonstrates that roughly a quarter of clear UCLs sampled  
358 in CSET had RH lower than 50%. Since PBL RH rarely falls below 60%, such layers are UCLs  
359 in the lower FT. Therefore, UCLs are not solely a PBL phenomenon, although their frequency of  
360 occurrence in the lower FT appears to be lower than at the top of the PBL (Fig. 2).

361 The frequency of occurrence of UCLs occurring higher up in the FT is estimated from the  
362 CSET transit legs (altitude 6.5-7 km) that were conducted before and after the low-level sampling.  
363 Accounting for the lower density at the transit level to estimate the concentration that these layers  
364 would have when compressed adiabatically to 1.5 km altitude, we estimate UCL frequency when  
365 subsided to be approximately 15%, which is quite close to the observed frequency of clear UCLs  
366 in the lower FT (2-3 km altitude) shown in Fig. 2. Thus, it seems reasonable to conclude that

367 the subsiding FT has the potential to be a contributor to occurrences of UCLs found in the PBL.  
368 However, the fraction of cloudy samples classified as UCLs is much higher (30-40%) in the upper  
369 PBL (Fig. 2). Given this and the tendency for clear UCLs below 3km altitude to be associated with  
370 high RH, it seems likely that many clear UCLs in the PBL originate from the evaporation of UCL  
371 clouds in the PBL. In part II of this paper, we use a microphysical model to explore a mechanism  
372 for the production of UCL clouds, and conclude that collision-coalescence in cumulus updrafts is  
373 likely sufficient to explain the observed UCLs in the PBL.

374 Mean liquid water contents in UCL clouds tend to be relatively low compared with typical  
375 marine stratocumulus clouds, and the majority of the condensate tends to consist of droplets with  
376 radii exceeding  $37.5 \mu\text{m}$  that are measured by the 2DC probe (Fig. 4(c)). The total condensate  
377 (approximately represented by the sum of the CDP and 2DC  $q_L$ ) does not have strong height  
378 dependence and is typically  $0.1\text{-}0.15 \text{ g m}^{-3}$ . We use the CDP and 2DC size distributions to  
379 estimate the mean effective radii ( $r_e$ ) of UCL clouds and find mean values of  $20\text{-}25 \mu\text{m}$  (Fig. 4(d)).  
380 As might be expected given that droplets in the 2DC size range constitute the majority of the  
381 condensate, the mean effective radii derived from only the CDP probe (droplets with radii smaller  
382 than  $25 \mu\text{m}$ ) are considerably smaller (see e.g., Wood, 2000), indicating that condensate in small  
383 drizzle drops contributes significantly to  $r_e$  in UCL clouds and therefore influences significantly  
384 the optical properties of UCL clouds.

### 385 *b. Turbulence*

386 Clouds in UCLs tend to be much less turbulent than non-UCL clouds in the same altitude range.  
387 Fig. 6 shows estimates of inertial range turbulence derived every second using windowed (1 Hz)  
388 vertical velocity variance ( $\sigma_w^2$ ) from the high frequency (25 Hz) vertical wind measurements.  
389 Cloudy UCLs typically have mean  $\sigma_w^2$  from  $0.015\text{-}0.03 \text{ m}^2\text{s}^{-2}$ , whereas cloudy samples that are

390 non UCLs have values 3-5 times higher than this. This difference is consistent at all longitudes  
391 (not shown). Differences for clear samples are weaker, as are the turbulence levels themselves.  
392 Above 1 km altitude, clear UCLs have turbulence levels close to the likely minimum value that  
393 can be accurately determined with the aircraft turbulence probe (roughly  $0.005 \text{ m}^2 \text{ s}^{-2}$ ). In the  
394 altitude range 0.5-1.5 km, non UCL clear regions have turbulence levels roughly 30-100% higher  
395 than UCL clear regions, but above 1.5 km the differences are smaller. Thus, it may be concluded  
396 that both clear and cloudy UCLs, are typically significantly more quiescent than non-UCL equiv-  
397 alents at similar levels. There are likely some important consequences of low turbulence levels  
398 in UCLs. First, a UCL that mixes with air either above or below would typically result in an in-  
399 crease in aerosol concentration in the layer. Second, introduction of drier air via turbulent mixing  
400 into a saturated, cloudy UCL would also tend to result in cloud dessication. The low turbulence  
401 levels observed in UCLs, and the implied suppression of vertical mixing, therefore probably helps  
402 to explain their high frequency of occurrence. It may also be indicative of the process by which  
403 UCLs are formed in the first place. Theoretical and observational evidence suggests that Cu clouds  
404 detrain air preferentially at levels where there is static stability in the environmental profile (Ray-  
405 mond and Blyth, 1986; Bretherton and Smolarkiewicz, 1988; Perry and Hobbs, 1996; Zuidema,  
406 1998). If UCL clouds are predominantly formed in this way, then one would expect suppressed  
407 turbulence as the rising parcels in the cumulus encountered regions of increased buoyancy, caus-  
408 ing deceleration of vertical motion. Visual evidence of a rather laminar appearance of many of the  
409 layer clouds observed in CSET is presented in section 5.

## 410 **5. Case studies**

411 The statistical analysis of the properties of UCLs determined using *in situ* aircraft observations  
412 reveals that both cloudy and clear UCLs are common, especially in the upper parts of the marine

413 PBL above the surface mixed layer. In this section, we combine these observations with aircraft  
414 and satellite remote sensing information to provide additional context to highlight properties of  
415 UCL clouds that are difficult to ascertain using *in situ* data alone.

416 *a. 29 July 2015, Research Flight 11*

417 On 29 July 2015, CSET RF11 returned from Hilo to Sacramento and flew for 11 minutes (19:37-  
418 19:48 UTC) through the upper levels of an aggregated cumulus cluster situated at 29°N, 145°W  
419 approximately 1500 km to the northwest of Hawaii. Fig. 7(a) shows visible satellite imagery of  
420 the cluster indicating the track that the G-V flew through the system, while Fig. 7(b,c) show pho-  
421 tographs indicating extensive quasi-laminar layer clouds at a level similar to that of the aircraft  
422 ( $\sim 2750$  m). Just before exiting the cluster at 19:47 the G-V climbed slightly to an altitude of  
423 3350 m. The G-V passed through both clear and cloudy UCLs, as indicated by  $N_a$  and  $N_d$  con-  
424 centrations (green and dark blue dots respectively in Fig. 7(d)) below  $10 \text{ cm}^{-3}$ . Unfortunately, the  
425 UHSAS instrument suffered overheating during the period shown, but total aerosol concentrations  
426  $N_{CN}$  sampled in clear layers from 19:39-19:47 UTC after the UHSAS overheated remained very  
427 low and in the range  $10\text{-}30 \text{ cm}^{-3}$  (red dots in Fig. 7(d)), indicating likely clear UCLs associated  
428 with the high relative humidity air. In the dry free tropospheric air prior to the aircraft overfly-  
429 ing the cluster (19:36-19:37) and after the overpass (19:47-19:48) there was a marked increase in  
430 aerosol concentrations, indicating that the source of the sampled UCLs in this case is associated  
431 with the humid air in the upper reaches of the cumulus cluster.

432 Lidar backscatter imagery from the HSRL below the aircraft (Fig. 7(e)) indicates the presence  
433 of numerous horizontally-extensive cloud layers with top altitudes ranging from 1.5-2.5 km. Per-  
434 haps half of the cloud layers with tops around 1.5-2.0 km attenuate the lidar, but many do not.  
435 Those layers above 2 km rarely attenuated the lidar despite some of them (e.g., those from 19:42-

436 19:44 UTC) being as thick as 1 km. We refer to these thin clouds overlying the cumulus cluster  
437 as *veil clouds*. The G-V passed through the tops of these optically thin cloud layers from 19:41-  
438 19:43 UTC and measured condensate consisting of both cloud drops and small drizzle-sized drops  
439 (measured with the CDP and 2DC respectively). Total condensate amounts ranged from 0.04-  
440  $0.1 \text{ g m}^{-3}$  (Fig. 7(d)). Measured  $N_d$  in these UCL clouds was typically  $1\text{-}3 \text{ cm}^{-3}$  with effective  
441 radii ranging from  $20\text{-}50 \mu\text{m}$  (cyan dots in Fig. 7(d)). Despite relatively low condensate loadings,  
442 most of the UCL clouds detected by the lidar here produced detectable radar echoes with reflectiv-  
443 ities in the range  $-15$  to  $-25$  dBZ (Fig. 7(f)), consistent with significant condensate loading in the  
444 form of drizzle drops. Thus, in this case and in many others not shown here, the optically thin veil  
445 clouds are mostly existing in UCLs.

446 A remarkable feature of the optically thin veil clouds sampled in this cluster, given that they  
447 contain appreciable drizzle-sized drops, is that many have very sharply defined bases (seen in the  
448 HSRL and HCR images and also in the photographs in Fig. 7) with little evidence of virga in  
449 most cases, other than in some of the most geometrically thick but optically thin layer clouds (e.g.,  
450 19:43-19:44). This suggests that many of the veil clouds may not simply be decaying remnants of  
451 precipitating layers but may be being actively maintained. Evidence for mesoscale lofting of PBL  
452 air within aggregated trade cumulus clusters is found in large eddy simulations (Bretherton and  
453 Blossey, 2017), where such vertical motion appears to be supported by circulations that develop  
454 between the suppressed regions surrounding clusters and the active clusters themselves. How  
455 such mesoscale ascent, and how the UCL cloud microphysics might each impact cloud cover and  
456 longevity of veil clouds are questions worthy of further inquiry.

457 Although in this case we do not have an G-V profile to provide vertical context, a decoupled  
458 PBL structure is evident throughout the sampling period. This is clearest before 19:37 and after  
459 19:44 UTC when there is a relatively strongly scattering surface mixed layer extending up to 1 km

460 overlain with layers of much lower scattering. In places (e.g., at 2.5 km altitude from 19:36-19:39,  
461 and just below 1.5 km altitude between 19:44-19:45 UTC) there are curious horizontally-extensive  
462 layers of increased lidar scattering. It is possible that these layers are caused by the evaporation of  
463 cloud droplets (the scattering layers are horizontally adjacent to cloud layers).

464 In addition to the qualitative information provided by the lidar and radar sensors on the G-V, a  
465 real strength of these instruments is in their combination with each other to produce quantitative  
466 retrievals of cloud properties. The radar-lidar sensor synergy is at its best for the optically thin  
467 veil clouds, where lidar backscatter is available throughout the cloud layer. Fig. 8 shows retrievals  
468 of liquid water content and droplet effective radius using a variant of the O'Connor et al. (2005)  
469 method (Schwartz and Ghate, 2017), for the same 29 July cloud case shown in Fig. 7. Retrieved  
470 liquid water contents range up to  $0.5 \text{ g m}^{-3}$  are found, especially lower down in the cloud layer  
471 below the level of the G-V. Effective radii exceeding  $20 \mu\text{m}$  are found in much of the cloud layer,  
472 yielding estimates of cloud droplet concentration  $N_d$  that are a few  $\text{cm}^{-3}$ , and in good agreement  
473 with *in situ* estimates (Fig. 8d). In this study, we show only this one case, but future work could  
474 include a systematic analysis of the lidar-radar retrievals from all the CSET flights to help quantify  
475 their frequency, liquid water path and droplet concentration.

#### 476 *b. 27 July 2015, Research Flight 10*

477 On 27 July 2015, CSET RF10 flew from Sacramento to Hilo and sampled large clusters of aggre-  
478 gated cumulus convection and associated veil clouds west of the Sc-Cu transition ( $138^\circ\text{W}$ ). A time  
479 lapse set of GOES visible imagery illustrates the temporal evolution of optically thin veil clouds  
480 (reflectances of  $\sim 0.2$ ) and demonstrates their relationship with bright (reflectances exceeding 0.4)  
481 Cu clouds (Fig. 9). In this case, the G-V flew a straight leg through a stratiform cloud at the top of  
482 the PBL ( $\sim 2300 \text{ m}$ ). Low cloud droplet concentrations measured with the G-V were in the UCL

483 range during the latter half of the leg (Fig. 10c). Effective radii exceeds 20  $\mu\text{m}$  over much of this  
484 leg (Fig. 10), even reaching as high as 40-60  $\mu\text{m}$ . MODIS-retrieved cloud droplet effective radii  
485 are also very high over a large fraction of the cloud system (Fig. 10b) with independent estimates  
486 from the combined CDP/2D-C and HOLODEC in excellent agreement (Fig. 10d).

487 The temporal evolution seen in the GOES imagery is particularly revealing, as it demonstrates  
488 several key aspects the lifecycle of veil clouds. Active Cu clouds appear to be necessary for  
489 the initial formation of the stratiform veil clouds (Cu without veil clouds develop them after 1-  
490 3 hours as seen in Fig. 9e-h). After the active Cu weaken and dissipate (e.g., Fig. 9e-g), the  
491 veil clouds persist for a few more hours but then eventually evaporate (Fig. 9h,i). In this case  
492 there was relatively little vertical shear of the horizontal wind within the PBL, and the veil cloud  
493 locations remain quite close to the Cu that produce them. However, in other cases with directional  
494 wind shear in the PBL, veil clouds in the upper PBL can move in a direction opposite to the  
495 direction of motion of the Cu, and this sometimes leaves isolated patches of optically thin veil  
496 clouds completely disconnected from the Cu that generated them in the first place. Fig. 9g also  
497 demonstrates that aggregated Cu clusters are associated with rather large increases in column  
498 water vapor (from 25 mm in the suppressed regions to as high as 40 mm in the Cu patches). This  
499 concentration of moisture in the regions of active Cu appears to be a typical feature of shallow  
500 convective self-aggregation, as indicated by large domain large eddy simulations (Bretherton and  
501 Blossey, 2017).

502 *c. 19 July 2015, Research Flight 7*

503 The failure rate for visible/near infrared (VIS/NIR) cloud optical property retrievals from  
504 NASA's Moderate Resolution Imaging Spectroradiometer (MODIS) is approximately 10% Cho  
505 et al. (2015). The most prevalent type of failure of, explaining 60-85% of all failed retrievals, is

506 that the cloud effective radius implied by the combination of visible and NIR channels is larger  
507 than the upper bound ( $30 \mu\text{m}$ ) in the look-up table Cho et al. (2015). The mean UCL cloud effec-  
508 tive radus determined using *in situ* cloud probe data is  $\sim 20 \mu\text{m}$  (Fig. 4d), but that effective radii  
509 frequently approach and sometimes exceed  $30 \mu\text{m}$  (see part II of this paper, O et al. (2017)). Given  
510 this, the failure of MODIS retrievals due the presence of large droplets should not be surprising.

511 Here, we examine data from CSET Research Flight 7 on 19 July 2015, and compare *in situ* and  
512 satellite retrieved cloud properties, in this case from the GOES-15 geostationary satellite. Photo-  
513 graphic and satellite observations show intermingled shallow cumulus convection and stratiform  
514 veil clouds during the G-V traverse through an extensive field of open cellular clouds on a return  
515 leg from Hawaii to Sacramento (Fig. 11a-c). The aircraft remained within clouds with similar  
516 open cellular morphology from  $152^\circ\text{W}$  to  $135^\circ\text{W}$ , a distance of over 1700 km sampled over ap-  
517 proximately four hours (18:10-22:15 UTC). During this period the aircraft made numerous cloud  
518 penetrations, and GOES-15 cloud product retrievals (using a combination of  $0.63$  and  $3.9 \mu\text{m}$   
519 wavelength channels, Minnis et al., 2008a,b, 2011) were processed for the 17 images (15 minute  
520 temporal resolution) obtained during the aircraft sampling period. Retrievals from all pixels within  
521 a  $2^\circ$  swath around the aircraft track were used to produce histograms of cloud droplet concentration  
522 (Painemal et al., 2015) and effective radius for the transect that are compared with those derived  
523 from the *in situ* data (Fig. 11d,e). Although there is a rather large scale mismatch between the  
524 aircraft *in situ* microphysical estimates and those from the rather coarse ( $\sim 4$  km) GOES pixels,  
525 the two estimates agree fairly well, with median droplet concentrations  $N_d$  of  $7 \text{ cm}^{-3}$  (GOES) and  
526  $5 \text{ cm}^{-3}$  (*in situ*), and median effective radii of  $24 \mu\text{m}$  (GOES) and  $21 \mu\text{m}$  (*in situ*). The histograms  
527 do show some differences, with larger tails seen in the *in situ* data, which may be symptomatic of  
528 the different spatial scales of the measurements, but in general the agreement between the aircraft  
529 and satellite retrievals are rather good. This is somewhat surprising because these open cellular

530 regions do contain a significant fraction of clouds that are not particularly well-suited to VIS/NIR  
531 satellite retrievals, which assume plane parallel clouds, and may suffer from ambiguity problems  
532 for clouds of low optical thickness. Despite this, these findings provide some encouragement that  
533 passive satellite retrievals in such aggregated cumulus clusters may not be as poor as has some-  
534 times been supposed, especially if the retrievals are used statistically to characterize the average  
535 microphysical properties of cloud systems rather than at the pixel level.

536 Examination of effective radius images from GOES-15 in the region (e.g., Fig. 11b) indicates  
537 that extensive regions of the northeastern Pacific are typically covered with clouds with effective  
538 radii of 25-30  $\mu\text{m}$  at any one time, a finding consistent with our observations documenting the  
539 frequent occurrence of UCL clouds in the region. Thus, it is likely that even if the quantitative  
540 accuracy of the GOES-15 retrievals is poor, the estimates may still serve as a useful means to  
541 identify and track cloudy ultra-clean layers. Further work detailing more extensive comparisons  
542 of aircraft and satellite data will be needed to ascertain the extent to which this is the case.

## 543 **6. Aircraft remote sensing of optically thin clouds**

544 As can be readily seen in Fig. 7, a rather high fraction of the clouds sampled in CSET do not fully  
545 attenuate the HSRL lidar beam. The HSRL-determined cloud mask (Schwartz and Ghate, 2017)  
546 identifies clouds detected by the HSRL and also provides a flag indicating whether the beam has  
547 been attenuated. We use all zenith-pointing data from the near-surface level legs (see section 2a)  
548 to ensure that HSRL has a view of the entire column and therefore can determine the low cloud  
549 cover. For each lidar profile, we first determine if low clouds are present using the cloud mask.  
550 Here, low clouds are defined as those with HSRL-detected bases below 3 km altitude. These data  
551 can be used to determine the overall low cloud cover. Then, profiles that contain low cloud are  
552 separated into three categories: (i) optically thick low clouds, which are those profiles containing

553 a single low cloud layer that fully attenuates the lidar beam; (ii) a single layer of optically thin low  
554 cloud, i.e., a low cloud that does not fully attenuate the lidar, determined by detecting a lidar signal  
555 from above the cloud layer; (iii) multiple layers of low clouds. This latter category could contain  
556 multiple optically thin low cloud layers, or optically thin low cloud layers below optically thick  
557 low cloud layer above. For the single layer optically thin cloud we are able to determine both the  
558 cloud layer base and the top, and thus the cloud thickness.

559 Cloud statistics from all CSET flights are aggregated into  $5^\circ$  longitude bins from  $125\text{-}155^\circ\text{W}$   
560 (Fig. 12). Consistent with existing satellite climatologies (e.g., Karlsson et al., 2010; Leahy et al.,  
561 2012), low cloud cover decreases moving westward as the MBL transitions from relatively shall-  
562 low at the location of the stratocumulus maximum around  $130\text{-}135^\circ\text{W}$  to a deeper MBL close to  
563 Hawaii. Low cloud cover decreases as the MBL deepens and the stratocumulus-cumulus transi-  
564 tion occurs. The fraction of low clouds that are optically thin increases westward from below 0.4  
565 east of  $135^\circ\text{W}$  to 0.6-0.7 west of  $140^\circ\text{W}$ . These high values are consistent with CALIPSO lidar  
566 based climatologies (Leahy et al., 2012; Guzman et al., 2017). We further subdivide the optically  
567 thin low clouds into those occurring in the upper MBL (bases at or above 1 km altitude) that  
568 we refer to as veil clouds ( $f_{low,thin,z>1km}$ ) and those with bases near the lifting condensation level  
569 (bases below 700 m altitude,  $f_{low,thin,z<700m}$ ). Fig. 12 shows that most of the optically thin low  
570 clouds observed in CSET occurred in the upper PBL above 1 km, and are therefore considered  
571 veil clouds. Optically thin LCL clouds are far less numerous and nowhere contribute more than  
572 15% of all optically thin low clouds. The increasing frequency of optically thin low clouds moving  
573 westwards mirrors the increase in frequency of both clear and cloudy UCLs (Fig. 3), consistent  
574 with other lines of evidence indicating a tight connection between strongly depleted layers and the  
575 occurrence of optically thin veil clouds.

576 The heights and thicknesses of all the optically thin cloud layers detected by the zenith-pointing  
577 HSRL during CSET are summarized in Fig. 13. The figure shows box-whisker plots showing  
578 geometrical thickness vs cloud base height, with the blue curve providing the cumulative frequency  
579 of cloud base height. The median cloud base height of the optically thin clouds is 1150 m, with  
580 10th and 90th percentiles of 625 m and 1700 m respectively. Mean geometrical thicknesses of the  
581 optically thin layers are just over 200 m (medians are a little lower,  $\sim 180$  m), with little evidence  
582 of a systematic variation of thickness as base height increases. Some optically thick layers are  
583 as thick as 300-400 m; the optically thin clouds shown in Fig. 7 exceed this thickness in places.  
584 Most optically thin clouds are therefore occurring in the upper PBL above the surface mixed layer,  
585 whose top is typically 600-700 m. The height distribution of the fraction of cloudy samples that  
586 are UCLs determined using *in situ* data (Fig. 2) also shows few UCLs in the surface mixed layer,  
587 and a rapid increase with height, which is consistent with the idea that many optically thin low  
588 clouds are also UCL clouds. The red dashed line in Fig. 13 shows the cumulative distribution of  
589 the heights of UCL clouds determined using the *in situ* G-V cloud probe data. It should be stressed,  
590 however, that the cumulative frequency distribution of the heights of UCL clouds may be biased  
591 by non uniform height sampling by the G-V. To attempt to mitigate this potential sampling bias,  
592 we normalized the number of cloudy UCL samples identified at each level by the total number of  
593 aircraft samples at each level. The cumulative distributions of the lidar-determined optically thin  
594 cloud base height and the *in situ* determined UCL cloud height are similar, providing additional  
595 evidence of a close connection between optically thin clouds and UCL clouds. The two cumulative  
596 distributions do diverge somewhat at heights above about 1200 m, with almost twice as many  
597 cloudy UCLs occurring above 1500 m than optically thin clouds.

## 598 7. Discussion

599 They two key findings in this paper are that (a) ultraclean layers are a frequent occurrence in  
600 the MBL in the stratocumulus to cumulus transition region of the northeastern subtropical Pacific  
601 Ocean; (b) many clouds in UCLs tend to be optically thin, quasi-laminar stratiform veil clouds  
602 containing large droplets but relatively low liquid water contents. These clouds tend to occur in  
603 the upper part of the decoupled trade wind PBL and are only infrequently found in the surface  
604 mixed layer below 600 m or so. Veil clouds are particularly common in cases of open mesoscale  
605 cellular convection and in large ( $\sim 100$  km wide) aggregated cumulus clusters. The tendency  
606 for a high frequency of optically thin clouds in open and disorganized cellular convection has  
607 been previously observed (see, e.g., the cloud optical thickness distributions shown in Fig. 14  
608 in Muhlbauer et al. (2014)) but their microphysical and detailed structural properties have not  
609 been previously described. Veil clouds tend to be geometrically as well as optically thin, and  
610 so their low optical thickness can be attributed to a combination of their thinness and their very  
611 low droplet concentrations (optical thickness scales with  $N_d^{\frac{1}{3}}$ ). Veil clouds are relatively free of  
612 turbulence, and so assume a laminar like structure when viewed from the side. In the heart of the  
613 SCT upwind of Hawaii, the aircraft observations suggest that  $\sim 50\%$  of cloudy columns contain  
614 UCL clouds, and a similar fraction of low clouds is optically thin in this region, again consistent  
615 with the idea that most veil clouds are severely depleted of droplets. This begs the question of  
616 how UCL/veil clouds are formed, how they are maintained, how they eventually dissipate and  
617 what aerosol they leave behind upon evaporation. Visible geostationary imagery over a period of  
618 several hours demonstrates that veil clouds tend to form 1-2 hours after the formation of bright,  
619 active Cu cloud clusters, and then persist for roughly the same amount of time once the active  
620 Cu weaken and disappear. Although we have not conducted an exhaustive study of veil cloud

lifetimes, the veil clouds examined here last for several hours in total. Thus, veil clouds are a common feature of aggregated cumulus patches, and likely constitute the majority of the cloud cover within these patches.

*a. What determines the lifecycle and microphysics of UCLs and veil clouds?*

In Part II of this paper (O et al., 2017) we use microphysical parcel modeling to explore the idea that UCL clouds are created by rapid removal of cloud droplets by collision-coalescence in the moist updrafts of the active Cu clouds in aggregated clusters or in open cells, so that when these Cu detrain saturated air, it is already strongly depleted in droplets. This detrainment mechanism for UCL cloud formation appears to capture some key microphysical properties of UCL clouds, and can explain their longitudinal variation within the stratocumulus to cumulus transition, but other hypotheses may also explain them. For example, if aggregated Cu clusters are regions of weak mesoscale ascent, as is suggested in large eddy simulations (e.g., Bretherton and Blossey, 2017) then it might be possible for UCLs to form *in situ* rather than from detrainment of condensate from the more active Cu towers. However, if this were the case, then it is difficult to explain how the accumulation mode aerosol concentrations  $N_a$  in clear UCLs can have formed. Slow ascent on the order of  $\sim 1 \text{ cm s}^{-1}$  (as found in the simulations of Bretherton and Blossey (2017) in the aggregated Cu clusters) would indeed tend to result in rather low  $N_a$  even when  $N_a$  exceeds the UCL threshold of  $10 \text{ cm}^{-3}$ . However, such a cloud would also contain a significant number of interstitial unactivated accumulation mode aerosol that would remain in place after the cloud dissipated. Thus, it is difficult to explain the frequent occurrence of clear UCLs with such a formation mechanism. Mesoscale ascent could, by providing condensate, enhance veil cloud longevity compared with a case with no mesoscale ascent, and may help to explain their observed lifetimes of several hours. The terminal velocity of cloud droplets increases for  $1.2 \text{ cm s}^{-1}$  for

644 droplets with radii of  $10\ \mu\text{m}$  to  $5\ \text{cm s}^{-1}$  for droplets with radii of  $20\ \mu\text{m}$ , indicating that sustained  
645 mesoscale ascent of several  $\text{cm s}^{-1}$  would be required to maintain a UCL cloud against desiccation  
646 by sedimentation of droplets out of the layer. Whether such mesoscale lofting occurs for a suf-  
647 ficiently long time within aggregated cumulus clusters warrants further study using large domain  
648 cloud resolving models. It is possible that size sorting of droplets occurs in UCL clouds over the  
649 cloud lifecycle, such that the largest droplets are removed by sedimentation out of the cloud base,  
650 leaving behind smaller droplets. This temporal evolution of UCL cloud microphysics is explored  
651 in Part II of this paper.

652 *b. Condensate loading and optical thickness of veil clouds*

653 Based on the measurements presented here, we can make a first attempt to quantify the mean  
654 condensate loading contained in veil clouds in the SCT. First, we note the mean veil cloud liquid  
655 water contents (sum of the CDP and 2DC liquid water contents) are  $\sim 0.12\ \text{g m}^{-3}$  (Fig. 4c). It  
656 is worth noting that an uncertain, but likely significant fraction of the condensate is contained in  
657 small drizzle-sized drops larger than the CDP upper limit ( $25\ \mu\text{m}$  radius) but smaller than the  
658 first reliable bin of the 2DC ( $37.5\ \mu\text{m}$  radius), so this condensate loading may be a lower limit.  
659 Next, we found that the mean geometrical thickness of veil clouds is just over  $200\ \text{m}$  (Fig. 13).  
660 The product of the mean veil cloud liquid water content and the geometrical thickness yields an  
661 estimate of the mean liquid water path (LWP) of approximately  $25\ \text{g m}^{-2}$ , which is significant,  
662 although it should be pointed out that this LWP is close to the detection limit for passive microwave  
663 retrievals (Wentz, 1997), making accurate spaceborne quantification of condensate loading in veil  
664 clouds challenging.

665 Cloud optical thickness (see e.g., George and Wood, 2010) is related to LWP and  $N_d$  via

$$\tau \approx KN_d^{\frac{1}{3}}LWP^{\frac{5}{6}} \quad (1)$$

666 where  $K$  is a weak function of temperature and pressure and  $K \sim 0.25 \text{ kg}^{-5/6} \text{ m}^3$  for  $T = 280 \text{ K}$  and  
 667  $p = 850 \text{ hPa}$ . For the typical veil cloud LWP of  $25 \text{ g m}^{-2}$  and  $N_d \sim 5 \text{ cm}^{-3}$  (Fig. 4a), yielding  $\tau \sim 2$ .  
 668 This low value of  $\tau$  is consistent with the frequent penetration of veil clouds by the lidar. If veil  
 669 clouds did not have such low  $N_d$ , but instead had values close to the accumulation model aerosol  
 670 in the surface mixed layer ( $\sim 75 \text{ cm}^{-3}$ , see part II of this paper), then from Eqn. 1,  $\tau \sim 5$ , in which  
 671 case lidar would not penetrate the cloud. Indeed, for a threshold optical depth of 3 below which  
 672 lidar will penetrate (Winker and Poole, 1995),  $N_d$  must be below approximately  $20 \text{ cm}^{-3}$ . Thus,  
 673 the low optical thickness of veil clouds is strongly contingent on their low droplet concentrations.  
 674 Given the frequent occurrence of veil clouds, this implies that there may be a direct control of  
 675 regional cloud albedo by precipitation through coalescence scavenging.

### 676 *c. On the radiative susceptibility of veil clouds*

677 CSET data show that many UCL clouds, despite being optically and geometrically thin, have  
 678 radar echoes significantly higher than  $-30 \text{ dBZ}$  such that they can be detected with the W-band  
 679 radar on the G-V aircraft and may even be detectable and characterizable using spaceborne radar.  
 680 Radar-lidar retrievals designed originally for quantification of the microphysical properties of driz-  
 681 zle show considerable promise for remote characterization of the microphysical properties, con-  
 682 densate loadings, and potentially the albedo susceptibility of veil clouds to changes in aerosol  
 683 loading. The relatively low albedo  $A$  of veil clouds does not drastically reduce their albedo suscep-  
 684 tibility to changes in droplet concentration  $N_d$ , which scales as  $dA/dN_d = A(1 - A)/3N_d$  (Platnick  
 685 and Twomey, 1994), and thus is only about one third lower for  $A = 0.2$  compared with  $A = 0.5$ .

686 On the other hand, the extremely low  $N_d$  in UCL clouds implies a very high albedo suscepti-  
687 bility (Twomey effect). Marine low clouds shoulder a large fraction of the global indirect aerosol  
688 forcing in climate models (e.g., Kooperman et al., 2012; Carslaw et al., 2013; Lee et al., 2016),  
689 so it is important to understand whether models have the appropriate albedo susceptibility given  
690 the tendency for large scale models to produce an insufficient fraction of optically thin clouds.  
691 That said, the albedo susceptibility defined by Platnick and Twomey (1994) does not deal with the  
692 extent to which  $N_d$  can be perturbed by increases in anthropogenic aerosol emissions. For MBL  
693 clouds over the remote oceans, the pathways by which  $N_d$  can be increased is either via increased  
694 marine shipping emissions (surface sources) or by the entrainment of long-range transported pol-  
695 luted air from the FT. But in order to perturb veil cloud  $N_d$ , surface air must first pass through  
696 cumulus clouds in order to reach the upper MBL. Coalescence scavenging in cumulus updrafts  
697 is likely to strongly limit the perturbation to veil cloud  $N_d$  that would be expected from a given  
698 increase in aerosol loading in the surface mixed layer. Thus, although veil clouds have high albedo  
699 susceptibility, microphysical processes driving their production may limit their ability to realize  
700 this susceptibility. Further, because veil clouds tend to be quiescent, their ability to entrain FT air  
701 and tap into the higher aerosol concentrations there may also be strongly limited too. Exploring  
702 the degree to which veil clouds can be perturbed by aerosol is an important topic for further study.

703 Beyond the Twomey effect, there is the question of whether the colloidal stability of low  $N_d$   
704 veil clouds may change if aerosol loadings increase. Answering this will require much greater  
705 understanding of the lifecycle of veil clouds and of the factors that control their coverage and  
706 lifetime. An additional aspect of the susceptibility of top of atmosphere radiation to changes in  
707 the optical thickness or coverage of veil clouds is that they likely have significant longwave cloud  
708 radiative effects by virtue of their temperature contrast with the surface and because emissivity is  
709 sensitive to droplet effective radius for typical UCL cloud droplet sizes (Wood, 2012).

710 *d. Outlook for future research*

711 Over the global oceans, roughly half of all low clouds do not fully attenuate spaceborne lidar  
712 Leahy et al. (2012), indicating the important contribution of optically thin MBL clouds to global  
713 cloud cover, especially in the subtropics and tropics. Less is known about the contribution of these  
714 clouds to the overall cloud radiative effect. In the SCT regions, these optically thin clouds com-  
715 prise a mixture of small fair weather trade cumulus clouds with bases typically tied to the lifting  
716 condensation level (LCL), and much more horizontally extensive stratiform veil-type clouds that  
717 reside in the upper MBL. Quantifying the contribution of each of these types of optically thin cloud  
718 to cloud cover and to overall radiative effect is a significant challenge, but one that can probably be  
719 largely achieved with current and planned active spaceborne remote sensing technology. Under-  
720 standing of the large-scale conditions that encourage/discourage UCL production is needed. As  
721 major a challenge is to accurately represent veil clouds and UCLs in large-scale numerical models  
722 and to understand how they may change in response to increasing greenhouse gases and to aerosol  
723 changes. Parameterizations that are able to treat cumulus convection and stratiform veil clouds as  
724 tightly coupled entities with strongly interacting microphysical and macrophysical processes are  
725 likely to bear most fruit.

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731 **References**

732 Ackerman, A. S., 2003: Enhancement of cloud cover and suppression of nocturnal driz-  
733 zle in stratocumulus polluted by haze. *Geophysical Research Letters*, **30** (7), doi:10.1029/  
734 2002GL016634, URL <http://doi.wiley.com/10.1029/2002GL016634>.

735 Albrecht, B., 2017: Cloud System Evolution in the Trades (CSET) Field Campaign.

736 Augstein, E., H. Riehl, F. Ostapoff, and V. Wagner, 1973: Mass and Energy Transports  
737 in an Undisturbed Atlantic Trade-Wind Flow. *Mon. Wea. Rev.*, **101** (2), 101–111, doi:  
738 10.1175/1520-0493(1973)101<0101:MAETIA>2.3.CO;2, URL [http://journals.ametsoc.  
739 org/offcampus.lib.washington.edu/doi/10.1175/1520-0493%281973%29101%3C0101%  
740 3AMAETIA%3E2.3.CO%3B2](http://journals.ametsoc.org/offcampus.lib.washington.edu/doi/10.1175/1520-0493%281973%29101%3C0101%3AMAETIA%3E2.3.CO%3B2).

741 Berner, A. H., C. S. Bretherton, R. Wood, and A. Muhlbauer, 2013: Marine boundary layer  
742 cloud regimes and POC formation in a CRM coupled to a bulk aerosol scheme. *Atmospheric  
743 Chemistry and Physics*, **13** (24), 12 549–12 572, doi:10.5194/acp-13-12549-2013, URL [http:  
744 //www.atmos-chem-phys.net/13/12549/2013/](http://www.atmos-chem-phys.net/13/12549/2013/).

745 Bodas-Salcedo, A., and Coauthors, 2014: Origins of the Solar Radiation Biases over the Southern  
746 Ocean in CFMIP2 Models\*. *Journal of Climate*, **27** (1), 41–56, doi:10.1175/JCLI-D-13-00169.  
747 1, URL <http://journals.ametsoc.org/doi/abs/10.1175/JCLI-D-13-00169.1>.

748 Bretherton, C. S., P. Austin, and S. T. Siems, 1995: Cloudiness and Marine Boundary Layer  
749 Dynamics in the ASTEX Lagrangian Experiments. Part II: Cloudiness, Drizzle, Surface Fluxes,  
750 and Entrainment. *J. Atmos. Sci.*, **52** (16), 2724–2735, doi:10.1175/1520-0469(1995)052<2724:  
751 CAMBLD>2.0.CO;2, URL [http://journals.ametsoc.org/doi/abs/10.1175/1520-0469%281995%  
752 29052%3C2724%3ACAMBLD%3E2.0.CO%3B2](http://journals.ametsoc.org/doi/abs/10.1175/1520-0469%281995%29052%3C2724%3ACAMBLD%3E2.0.CO%3B2).

753 Bretherton, C. S., and P. Blossey, 2017: Understanding mesoscale aggregation of shallow cumulus  
754 convection using large-eddy simulation.

755 Bretherton, C. S., and R. Pincus, 1995: Cloudiness and Marine Boundary Layer Dynamics  
756 in the ASTEX Lagrangian Experiments. Part I: Synoptic Setting and Vertical Structure. *J.*  
757 *Atmos. Sci.*, **52** (16), 2707–2723, doi:10.1175/1520-0469(1995)052<2707:CAMBLD>2.0.CO;  
758 2, URL [http://journals.ametsoc.org/doi/abs/10.1175/1520-0469%281995%29052%3C2707%  
759 3ACAMBLD%3E2.0.CO%3B2](http://journals.ametsoc.org/doi/abs/10.1175/1520-0469%281995%29052%3C2707%3ACAMBLD%3E2.0.CO%3B2).

760 Bretherton, C. S., and P. Smolarkiewicz, 1988: Gravity Waves, Compensating Subsidi-  
761 dence and Detrainment around Cumulus Clouds. *J. Atmos. Sci.*, **46** (6), 740–759, doi:  
762 10.1175/1520-0469(1989)046<0740:GWCSAD>2.0.CO;2, URL [http://journals.ametsoc.org.  
763 offcampus.lib.washington.edu/doi/citedby/10.1175/1520-0469%281989%29046%3C0740%  
764 3AGWCSAD%3E2.0.CO%3B2](http://journals.ametsoc.org.offcampus.lib.washington.edu/doi/citedby/10.1175/1520-0469%281989%29046%3C0740%3AGWCSAD%3E2.0.CO%3B2).

765 Bretherton, C. S., and M. C. Wyant, 1997: Moisture transport, lower-tropospheric sta-  
766 bility, and decoupling of cloud-topped boundary layers. *Journal of the atmospheric sci-*  
767 *ences*, **54** (1), 148–167, URL [http://journals.ametsoc.org/doi/full/10.1175/1520-0469\(1997\)  
768 054%3C0148:MTL TSA%3E2.0.CO%3B2](http://journals.ametsoc.org/doi/full/10.1175/1520-0469(1997)054%3C0148:MTL TSA%3E2.0.CO%3B2).

769 Carslaw, K. S., and Coauthors, 2013: Large contribution of natural aerosols to uncertainty in  
770 indirect forcing. *Nature*, **503** (7474), 67–71, doi:10.1038/nature12674, URL [http://www.nature.  
771 com/doi/finder/10.1038/nature12674](http://www.nature.com/doi/finder/10.1038/nature12674).

772 Cho, H.-M., and Coauthors, 2015: Frequency and causes of failed MODIS cloud prop-  
773 erty retrievals for liquid phase clouds over global oceans: FAILED MODIS CLOUD RE-  
774 TRIEVALS. *Journal of Geophysical Research: Atmospheres*, **120** (9), 4132–4154, doi:10.1002/  
775 2015JD023161, URL <http://doi.wiley.com/10.1002/2015JD023161>.

776 Engstrm, A., F. a.-M. Bender, R. J. Charlson, and R. Wood, 2015: The nonlinear relation-  
777 ship between albedo and cloud fraction on near-global, monthly mean scale in observations  
778 and in the CMIP5 model ensemble. *Geophys. Res. Lett.*, **42** (21), 2015GL066275, doi:10.  
779 1002/2015GL066275, URL [http://onlinelibrary.wiley.com/offcampus.lib.washington.edu/doi/](http://onlinelibrary.wiley.com/offcampus.lib.washington.edu/doi/10.1002/2015GL066275/abstract)  
780 [10.1002/2015GL066275/abstract](http://onlinelibrary.wiley.com/offcampus.lib.washington.edu/doi/10.1002/2015GL066275/abstract).

781 Field, P. R., R. J. Cotton, K. McBeath, A. P. Lock, S. Webster, and R. P. Allan, 2014: Improving  
782 a convection-permitting model simulation of a cold air outbreak: Simulation of a Cold Air  
783 Outbreak. *Quarterly Journal of the Royal Meteorological Society*, **140** (678), 124–138, doi:  
784 [10.1002/qj.2116](http://doi.wiley.com/10.1002/qj.2116), URL <http://doi.wiley.com/10.1002/qj.2116>.

785 Fugal, J. P., and R. A. Shaw, 2009: Cloud particle size distributions measured with an air-  
786 borne digital in-line holographic instrument. *Atmos. Meas. Tech.*, **2** (1), 259–271, doi:10.5194/  
787 [amt-2-259-2009](http://www.atmos-meas-tech.net/2/259/2009/), URL <http://www.atmos-meas-tech.net/2/259/2009/>.

788 George, R. C., and R. Wood, 2010: Subseasonal variability of low cloud radiative properties  
789 over the southeast Pacific Ocean. *Atmospheric Chemistry and Physics*, **10** (8), 4047–4063, doi:  
790 [10.5194/acp-10-4047-2010](http://www.atmos-chem-phys.net/10/4047/2010/), URL <http://www.atmos-chem-phys.net/10/4047/2010/>.

791 Grund, C. J., and E. W. Eloranta, 1991: University of Wisconsin High Spectral Resolution Lidar.  
792 *Opt. Eng.*, **30** (1), 6–12, doi:10.1117/12.55766, URL <http://dx.doi.org/10.1117/12.55766>.

793 Guzman, R., and Coauthors, 2017: Direct atmosphere opacity observations from CALIPSO pro-  
794 vide new constraints on cloud-radiation interactions: GOCCP v3.0 OPAQ Algorithm. *Journal of*  
795 *Geophysical Research: Atmospheres*, doi:10.1002/2016JD025946, URL [http://doi.wiley.com/](http://doi.wiley.com/10.1002/2016JD025946)  
796 [10.1002/2016JD025946](http://doi.wiley.com/10.1002/2016JD025946).

797 Hindman, E. E., W. M. Porch, J. G. Hudson, and P. A. Durkee, 1994: Ship-produced cloud lines  
798 of 13 July 1991. *Atmospheric Environment*, **28 (20)**, 3393–3403, doi:10.1016/1352-2310(94)  
799 00171-G, URL <http://www.sciencedirect.com/science/article/pii/135223109400171G>.

800 Hobbs, P. V., 2003: Clean air slots amid dense atmospheric pollution in southern Africa. *J. Geo-*  
801 *phys. Res.*, **108 (D13)**, 8490, doi:10.1029/2002JD002156, URL [http://onlinelibrary.wiley.com/](http://onlinelibrary.wiley.com/doi/10.1029/2002JD002156/abstract)  
802 [doi/10.1029/2002JD002156/abstract](http://onlinelibrary.wiley.com/doi/10.1029/2002JD002156/abstract).

803 Hudson, J. G., and Y. Xie, 1999: Vertical distributions of cloud condensation nuclei spectra  
804 over the summertime northeast Pacific and Atlantic Oceans. *J. Geophys. Res.*, **104 (D23)**,  
805 30 219–30 229, doi:10.1029/1999JD900413, URL [http://onlinelibrary.wiley.com.offcampus.lib.](http://onlinelibrary.wiley.com.offcampus.lib.washington.edu/doi/10.1029/1999JD900413/abstract)  
806 [washington.edu/doi/10.1029/1999JD900413/abstract](http://onlinelibrary.wiley.com.offcampus.lib.washington.edu/doi/10.1029/1999JD900413/abstract).

807 Karlsson, J., G. Svensson, S. Cardoso, J. Teixeira, and S. Paradise, 2010: Subtropical Cloud-  
808 Regime Transitions: Boundary Layer Depth and Cloud-Top Height Evolution in Models and  
809 Observations. *J. Appl. Meteor. Climatol.*, **49 (9)**, 1845–1858, doi:10.1175/2010JAMC2338.1,  
810 URL <http://journals.ametsoc.org/doi/abs/10.1175/2010JAMC2338.1>.

811 Kazil, J., H. Wang, G. Feingold, A. D. Clarke, J. R. Snider, and A. R. Bandy, 2011: Modeling  
812 chemical and aerosol processes in the transition from closed to open cells during VOCALS-  
813 REx. *Atmos. Chem. Phys.*, **11 (15)**, 7491–7514, doi:10.5194/acp-11-7491-2011, URL [http://](http://www.atmos-chem-phys.net/11/7491/2011/)  
814 [www.atmos-chem-phys.net/11/7491/2011/](http://www.atmos-chem-phys.net/11/7491/2011/).

815 Kooperman, G. J., M. S. Pritchard, S. J. Ghan, M. Wang, R. C. J. Somerville, and L. M. Russell,  
816 2012: Constraining the influence of natural variability to improve estimates of global aerosol  
817 indirect effects in a nudged version of the Community Atmosphere Model 5. *Journal of Geo-*  
818 *physical Research*, **117 (D23)**, doi:10.1029/2012JD018588, URL [http://doi.wiley.com/10.1029/](http://doi.wiley.com/10.1029/2012JD018588)  
819 [2012JD018588](http://doi.wiley.com/10.1029/2012JD018588).

820 Krueger, S. K., G. T. McLean, and Q. Fu, 1995: Numerical Simulation of the Stratus-to-  
821 Cumulus Transition in the Subtropical Marine Boundary Layer. Part II: Boundary-Layer  
822 Circulation. *J. Atmos. Sci.*, **52** (16), 2851–2868, doi:10.1175/1520-0469(1995)052<2851:  
823 NSOTST>2.0.CO;2, URL [http://journals.ametsoc.org.offcampus.lib.washington.edu/doi/abs/  
824 10.1175/1520-0469%281995%29052%3C2851%3ANSOTST%3E2.0.CO%3B2](http://journals.ametsoc.org.offcampus.lib.washington.edu/doi/abs/10.1175/1520-0469%281995%29052%3C2851%3ANSOTST%3E2.0.CO%3B2).

825 Leahy, L. V., R. Wood, R. J. Charlson, C. A. Hostetler, R. R. Rogers, M. A. Vaughan, and D. M.  
826 Winker, 2012: On the nature and extent of optically thin marine low clouds: OPTICALLY  
827 THIN MARINE LOW CLOUDS. *Journal of Geophysical Research: Atmospheres*, **117** (D22),  
828 n/a–n/a, doi:10.1029/2012JD017929, URL <http://doi.wiley.com/10.1029/2012JD017929>.

829 Lee, L. A., C. L. Reddington, and K. S. Carslaw, 2016: On the relationship between aerosol model  
830 uncertainty and radiative forcing uncertainty. *PNAS*, **113** (21), 5820–5827, doi:10.1073/pnas.  
831 1507050113, URL <http://www.pnas.org/content/113/21/5820>.

832 Lu, M.-L., W. C. Conant, H. H. Jonsson, V. Varutbangkul, R. C. Flagan, and J. H. Seinfeld, 2007:  
833 The Marine Stratus/Stratocumulus Experiment (MASE): Aerosol-cloud relationships in marine  
834 stratocumulus. *Journal of Geophysical Research*, **112** (D10), doi:10.1029/2006JD007985, URL  
835 <http://doi.wiley.com/10.1029/2006JD007985>.

836 Martin, G. M., D. W. Johnson, D. P. Rogers, P. R. Jonas, P. Minnis, and D. A. Hegg, 1995:  
837 Observations of the Interaction between Cumulus Clouds and Warm Stratocumulus Clouds in  
838 the Marine Boundary Layer during ASTEX. *J. Atmos. Sci.*, **52** (16), 2902–2922, doi:10.1175/  
839 1520-0469(1995)052<2902:OOTIBC>2.0.CO;2, URL [http://journals.ametsoc.org.offcampus.  
840 lib.washington.edu/doi/abs/10.1175/1520-0469\(1995\)052%3C2902:OOTIBC%3E2.0.CO;2](http://journals.ametsoc.org.offcampus.lib.washington.edu/doi/abs/10.1175/1520-0469(1995)052%3C2902:OOTIBC%3E2.0.CO;2).

841 Martin, G. M., D. W. Johnson, and A. Spice, 1994: The Measurement and Pa-  
842 rameterization of Effective Radius of Droplets in Warm Stratocumulus Clouds. *J. At-*

843 *mos. Sci.*, **51** (13), 1823–1842, doi:10.1175/1520-0469(1994)051<1823:TMAPOE>2.0.CO;  
844 2, URL [http://journals.ametsoc.org/doi/abs/10.1175/1520-0469%281994%29051%3C1823%  
3ATMAPOE%3E2.0.CO%3B2](http://journals.ametsoc.org/doi/abs/10.1175/1520-0469%281994%29051%3C1823%<br/>845 3ATMAPOE%3E2.0.CO%3B2).

846 Mauritsen, T., and Coauthors, 2011: An Arctic CCN-limited cloud-aerosol regime. *Atmospheric*  
847 *Chemistry and Physics*, **11** (1), 165–173, doi:10.5194/acp-11-165-2011, URL [http://www.  
atmos-chem-phys.net/11/165/2011/](http://www.<br/>848 atmos-chem-phys.net/11/165/2011/).

849 Miles, N. L., J. Verlinde, and E. E. Clothiaux, 2000: Cloud droplet size distributions in low-level  
850 stratiform clouds. *Journal of the atmospheric sciences*, **57** (2), 295–311, URL [http://journals.  
ametsoc.org/doi/abs/10.1175/1520-0469\(2000\)057%3C0295:CDSDIL%3E2.0.CO%3B2](http://journals.<br/>851 ametsoc.org/doi/abs/10.1175/1520-0469(2000)057%3C0295:CDSDIL%3E2.0.CO%3B2).

852 Miller, M. A., and B. A. Albrecht, 1995: Surface-Based Observations of Mesoscale Cu-  
853 mulusStratocumulus Interaction during ASTEX. *J. Atmos. Sci.*, **52** (16), 2809–2826,  
854 doi:10.1175/1520-0469(1995)052<2809:SBOOMC>2.0.CO;2, URL [http://journals.ametsoc.  
org.offcampus.lib.washington.edu/doi/abs/10.1175/1520-0469%281995%29052%3C2809%  
3ASBOOMC%3E2.0.CO%3B2](http://journals.ametsoc.<br/>855 org.offcampus.lib.washington.edu/doi/abs/10.1175/1520-0469%281995%29052%3C2809%<br/>856 3ASBOOMC%3E2.0.CO%3B2).

857 Minnis, P., and Coauthors, 2008a: Cloud Detection in Nonpolar Regions for CERES Using  
858 TRMM VIRS and Terra and Aqua MODIS Data. *IEEE Transactions on Geoscience and Re-  
859 mote Sensing*, **46** (11), 3857–3884, doi:10.1109/TGRS.2008.2001351.

860 Minnis, P., and Coauthors, 2008b: Near-real time cloud retrievals from operational and research  
861 meteorological satellites. Vol. 7107, 710 703–710 703–8, doi:10.1117/12.800344, URL [http://  
dx.doi.org/10.1117/12.800344](http://<br/>862 dx.doi.org/10.1117/12.800344).

863 Minnis, P., and Coauthors, 2011: CERES Edition-2 Cloud Property Retrievals Using TRMM  
864 VIRS and Terra and Aqua MODIS Data #x2014;Part I: Algorithms. *IEEE Transactions on Geo-*

- 865 *science and Remote Sensing*, **49** (11), 4374–4400, doi:10.1109/TGRS.2011.2144601.
- 866 Morcrette, J.-J., and Y. Fouquart, 1986: The Overlapping of Cloud Layers in Shortwave Ra-  
867 diation Parameterizations. *J. Atmos. Sci.*, **43** (4), 321–328, doi:10.1175/1520-0469(1986)  
868 043(0321:TOOCLI)2.0.CO;2, URL [http://journals.ametsoc.org/doi/abs/10.1175/1520-0469%](http://journals.ametsoc.org/doi/abs/10.1175/1520-0469%281986%29043%3C0321%3ATOOCCLI%3E2.0.CO%3B2)  
869 [281986%29043%3C0321%3ATOOCCLI%3E2.0.CO%3B2](http://journals.ametsoc.org/doi/abs/10.1175/1520-0469%281986%29043%3C0321%3ATOOCCLI%3E2.0.CO%3B2).
- 870 Muhlbauer, A., I. L. McCoy, and R. Wood, 2014: Climatology of stratocumulus cloud mor-  
871 phologies: microphysical properties and radiative effects. *Atmospheric Chemistry and Physics*,  
872 **14** (13), 6695–6716, doi:10.5194/acp-14-6695-2014, URL [http://www.atmos-chem-phys.net/](http://www.atmos-chem-phys.net/14/6695/2014/)  
873 [14/6695/2014/](http://www.atmos-chem-phys.net/14/6695/2014/).
- 874 Nam, C., S. Bony, J.-L. Dufresne, and H. Chepfer, 2012: The too few, too bright tropical low-cloud  
875 problem in CMIP5 models. *Geophys. Res. Lett.*, **39** (21), L21 801, doi:10.1029/2012GL053421,  
876 URL <http://onlinelibrary.wiley.com/doi/10.1029/2012GL053421/abstract>.
- 877 O, K.-T., R. Wood, and C. S. Bretherton, 2017: Ultraclean layers and optically thin clouds in the  
878 stratocumulus to cumulus transition. Part II: Depletion of cloud droplets and cloud condensation  
879 nuclei through collision-coalescence in marine boundary layer clouds. *J. Atmos. Sci.*, submitted.
- 880 O’Connor, E. J., R. J. Hogan, and A. J. Illingworth, 2005: Retrieving stratocumulus drizzle pa-  
881 rameters using Doppler radar and lidar. *Journal of Applied Meteorology*, **44** (1), 14–27, URL  
882 <http://journals.ametsoc.org/doi/abs/10.1175/JAM-2181.1>.
- 883 Painemal, D., P. Minnis, and M. Nordeen, 2015: Aerosol variability, synoptic-scale processes,  
884 and their link to the cloud microphysics over the northeast Pacific during MAGIC: NE PA-  
885 CIFIC CCN AND CLOUD MICROPHYSICS. *Journal of Geophysical Research: Atmo-*

886 *spheres*, **120** (10), 5122–5139, doi:10.1002/2015JD023175, URL [http://doi.wiley.com/10.1002/](http://doi.wiley.com/10.1002/2015JD023175)  
887 2015JD023175.

888 Perry, K., and P. V. Hobbs, 1996: Influences of Isolated Cumulus Clouds on the Humidity of  
889 Their Surroundings. *J. Atmos. Sci.*, **53** (1), 159–174, doi:10.1175/1520-0469(1996)053<0159:  
890 IOICCO>2.0.CO;2, URL [http://journals.ametsoc.org/doi/abs/10.1175/1520-0469\(1996\)053%](http://journals.ametsoc.org/doi/abs/10.1175/1520-0469(1996)053%3C0159:IOICCO%3E2.0.CO;2)  
891 3C0159:IOICCO%3E2.0.CO;2.

892 Petters, M. D., J. R. Snider, B. Stevens, G. Vali, I. Faloon, and L. M. Russell, 2006:  
893 Accumulation mode aerosol, pockets of open cells, and particle nucleation in the remote  
894 subtropical Pacific marine boundary layer. *J. Geophys. Res.*, **111** (D2), D02206, doi:10.  
895 1029/2004JD005694, URL [http://onlinelibrary.wiley.com/offcampus.lib.washington.edu/doi/](http://onlinelibrary.wiley.com/offcampus.lib.washington.edu/doi/10.1029/2004JD005694/abstract)  
896 10.1029/2004JD005694/abstract.

897 Platnick, S., and S. Twomey, 1994: Determining the Susceptibility of Cloud Albedo to Changes in  
898 Droplet Concentration with the Advanced Very High Resolution Radiometer. *J. Appl. Meteor.*,  
899 **33** (3), 334–347, doi:10.1175/1520-0450(1994)033<0334:DTSOCA>2.0.CO;2, URL [http://](http://journals.ametsoc.org/doi/abs/10.1175/1520-0450(1994)033%3C0334:DTSOCA%3E2.0.CO;2)  
900 [journals.ametsoc.org/doi/abs/10.1175/1520-0450\(1994\)033%3C0334:DTSOCA%3E2.0.CO;2.](http://journals.ametsoc.org/doi/abs/10.1175/1520-0450(1994)033%3C0334:DTSOCA%3E2.0.CO;2)

901 Raymond, D. J., and A. M. Blyth, 1986: A Stochastic Mixing Model for Non-  
902 precipitating Cumulus Clouds. *J. Atmos. Sci.*, **43** (22), 2708–2718, doi:10.1175/  
903 1520-0469(1986)043<2708:ASMMFN>2.0.CO;2, URL [http://journals.ametsoc.org/doi/abs/10.](http://journals.ametsoc.org/doi/abs/10.1175/1520-0469(1986)043%3C2708:ASMMFN%3E2.0.CO;2)  
904 1175/1520-0469(1986)043%3C2708:ASMMFN%3E2.0.CO;2.

905 Sandu, I., and B. Stevens, 2011: On the Factors Modulating the Stratocumulus to Cumulus Tran-  
906 sitions. *Journal of the Atmospheric Sciences*, **68** (9), 1865–1881, doi:10.1175/2011JAS3614.1,  
907 URL [http://journals.ametsoc.org/doi/abs/10.1175/2011JAS3614.1.](http://journals.ametsoc.org/doi/abs/10.1175/2011JAS3614.1)

- 908 Sandu, I., B. Stevens, and R. Pincus, 2010: On the transitions in marine boundary layer cloudiness.  
909 *Atmospheric Chemistry and Physics*, **10** (5), 2377–2391, URL <http://www.atmos-chem-phys.net/10/2377/>.  
910
- 911 Schwartz, C., and V. Ghate, 2017: Retrievals of cloud and precipitation using airborne lidar-radar  
912 data from the northeastern Pacific. *Journal of Atmospheric and Oceanic Technology*.
- 913 Sharon, T. M., B. A. Albrecht, H. H. Jonsson, P. Minnis, M. M. Khaiyer, T. M. van Reken, J. Seinfeld,  
914 and R. Flagan, 2006: Aerosol and cloud microphysical characteristics of rifts and gradients  
915 in maritime stratocumulus clouds. *Journal of the atmospheric sciences*, **63** (3), 983–997, URL  
916 <http://journals.ametsoc.org/doi/abs/10.1175/JAS3667.1>.
- 917 Stevens, B., G. Vali, K. Comstock, R. Wood, M. C. Van Zanten, P. H. Austin, C. S. Brether-  
918 ton, and D. H. Lenschow, 2005: POCKETS OF OPEN CELLS AND DRIZZLE IN MARINE  
919 STRATOCUMULUS. *Bulletin of the American Meteorological Society*, **86** (1), 51–57, doi:  
920 [10.1175/BAMS-86-1-51](http://journals.ametsoc.org/doi/abs/10.1175/BAMS-86-1-51), URL <http://journals.ametsoc.org/doi/abs/10.1175/BAMS-86-1-51>.
- 921 Stevens, B., and Coauthors, 2001: Simulations of Trade Wind Cumuli under a Strong  
922 Inversion. *J. Atmos. Sci.*, **58** (14), 1870–1891, doi:10.1175/1520-0469(2001)058<1870:  
923 SOTWCU>2.0.CO;2, URL [http://journals.ametsoc.org.offcampus.lib.washington.edu/doi/abs/  
924 10.1175/1520-0469\(2001\)058%3C1870:SOTWCU%3E2.0.CO;2](http://journals.ametsoc.org.offcampus.lib.washington.edu/doi/abs/10.1175/1520-0469(2001)058%3C1870:SOTWCU%3E2.0.CO;2).
- 925 Teixeira, J., and Coauthors, 2011: Tropical and Subtropical Cloud Transitions in Weather and  
926 Climate Prediction Models: The GCSS/WGNE Pacific Cross-Section Intercomparison (GPCI).  
927 *J. Climate*, **24** (20), 5223–5256, doi:10.1175/2011JCLI3672.1, URL [http://journals.ametsoc.  
928 org.offcampus.lib.washington.edu/doi/abs/10.1175/2011JCLI3672.1](http://journals.ametsoc.org.offcampus.lib.washington.edu/doi/abs/10.1175/2011JCLI3672.1).

- 929 Terai, C. R., C. S. Bretherton, R. Wood, and G. Painter, 2014: Aircraft observations of  
930 aerosol, cloud, precipitation, and boundary layer properties in pockets of open cells over  
931 the southeast Pacific. *Atmospheric Chemistry and Physics*, **14** (15), 8071–8088, doi:10.5194/  
932 acp-14-8071-2014, URL <http://www.atmos-chem-phys.net/14/8071/2014/>.
- 933 Vivekanandan, J., and Coauthors, 2015: A wing pod-based millimeter wavelength airborne cloud  
934 radar. *Geoscientific Instrumentation, Methods and Data Systems*, **4** (2), 161–176, doi:10.5194/  
935 gi-4-161-2015, URL <http://www.geosci-instrum-method-data-syst.net/4/161/2015/>.
- 936 Wentz, F. J., 1997: A well-calibrated ocean algorithm for special sensor microwave / imager. *J.*  
937 *Geophys. Res.*, **102** (C4), 8703–8718, doi:10.1029/96JC01751, URL <http://onlinelibrary.wiley.com/offcampus.lib.washington.edu/doi/10.1029/96JC01751/abstract>.
- 938
- 939 Williams, K. D., and Coauthors, 2012: The Transpose-AMIP II Experiment and Its Applica-  
940 tion to the Understanding of Southern Ocean Cloud Biases in Climate Models. *J. Climate*,  
941 **26** (10), 3258–3274, doi:10.1175/JCLI-D-12-00429.1, URL <http://journals.ametsoc.org/doi/abs/10.1175/JCLI-D-12-00429.1>.
- 942
- 943 Winker, D. M., and L. R. Poole, 1995: Monte-Carlo calculations of cloud returns for ground-  
944 based and space-based LIDARS. *Appl. Phys. B*, **60** (4), 341–344, doi:10.1007/BF01082269,  
945 URL <https://link.springer.com/article/10.1007/BF01082269>.
- 946 Wood, R., 2000: Parametrization of the effect of drizzle upon the droplet effective radius in stra-  
947 tocumulus clouds. *Q.J.R. Meteorol. Soc.*, **126** (570), 3309–3324, doi:10.1002/qj.49712657015,  
948 URL <http://onlinelibrary.wiley.com/doi/10.1002/qj.49712657015/abstract>.
- 949 Wood, R., 2012: Stratocumulus Clouds. *Monthly Weather Review*, **140** (8), 2373–  
950 2423, doi:10.1175/MWR-D-11-00121.1, URL <http://journals.ametsoc.org/doi/abs/10.1175/>

951 MWR-D-11-00121.1.

952 Wood, R., C. S. Bretherton, D. Leon, A. D. Clarke, P. Zuidema, G. Allen, and H. Coe,  
953 2011: An aircraft case study of the spatial transition from closed to open mesoscale cel-  
954 lular convection over the Southeast Pacific. *Atmos. Chem. Phys.*, **11** (5), 2341–2370, doi:  
955 10.5194/acp-11-2341-2011, URL <http://www.atmos-chem-phys.net/11/2341/2011/>.

956 Wood, R., K. K. Comstock, C. S. Bretherton, C. Cornish, J. Tomlinson, D. R. Collins, and  
957 C. Fairall, 2008: Open cellular structure in marine stratocumulus sheets. *Journal of Geo-*  
958 *physical Research*, **113** (D12), doi:10.1029/2007JD009371, URL [http://doi.wiley.com/10.1029/](http://doi.wiley.com/10.1029/2007JD009371)  
959 [2007JD009371](http://doi.wiley.com/10.1029/2007JD009371).

960 Wood, R., D. Leon, M. Lebsock, J. Snider, and A. D. Clarke, 2012: Precipitation driving  
961 of droplet concentration variability in marine low clouds: PRECIPITATION DRIVING OF  
962 DROP CONC. *Journal of Geophysical Research: Atmospheres*, **117** (D19), n/a–n/a, doi:  
963 10.1029/2012JD018305, URL <http://doi.wiley.com/10.1029/2012JD018305>.

964 Wyant, M. C., C. S. Bretherton, H. A. Rand, and D. E. Stevens, 1997: Numerical simulations and  
965 a conceptual model of the stratocumulus to trade cumulus transition. *Journal of the atmospheric*  
966 *sciences*, **54** (1), 168–192, URL [http://journals.ametsoc.org/doi/abs/10.1175/1520-0469\(1997\)](http://journals.ametsoc.org/doi/abs/10.1175/1520-0469(1997)054%3C0168%3ANSAACM%3E2.0.CO%3B2)  
967 [054%3C0168%3ANSAACM%3E2.0.CO%3B2](http://journals.ametsoc.org/doi/abs/10.1175/1520-0469(1997)054%3C0168%3ANSAACM%3E2.0.CO%3B2).

968 Zhou, X., P. Kollias, and E. R. Lewis, 2015: Clouds, Precipitation, and Marine Boundary Layer  
969 Structure during the MAGIC Field Campaign. *J. Climate*, **28** (6), 2420–2442, doi:10.1175/  
970 JCLI-D-14-00320.1, URL <http://journals.ametsoc.org/doi/abs/10.1175/JCLI-D-14-00320.1>.

971 Zuidema, P., 1998: The 600800-mb Minimum in Tropical Cloudiness Observed during  
972 TOGA COARE. *J. Atmos. Sci.*, **55** (12), 2220–2228, doi:10.1175/1520-0469(1998)055<2220:

973 TMMITC)2.0.CO;2, URL [http://journals.ametsoc.org/doi/abs/10.1175/1520-0469%281998%](http://journals.ametsoc.org/doi/abs/10.1175/1520-0469%281998%29055%3C2220%3ATMMITC%3E2.0.CO%3B2)  
974 [29055%3C2220%3ATMMITC%3E2.0.CO%3B2](http://journals.ametsoc.org/doi/abs/10.1175/1520-0469%281998%29055%3C2220%3ATMMITC%3E2.0.CO%3B2).

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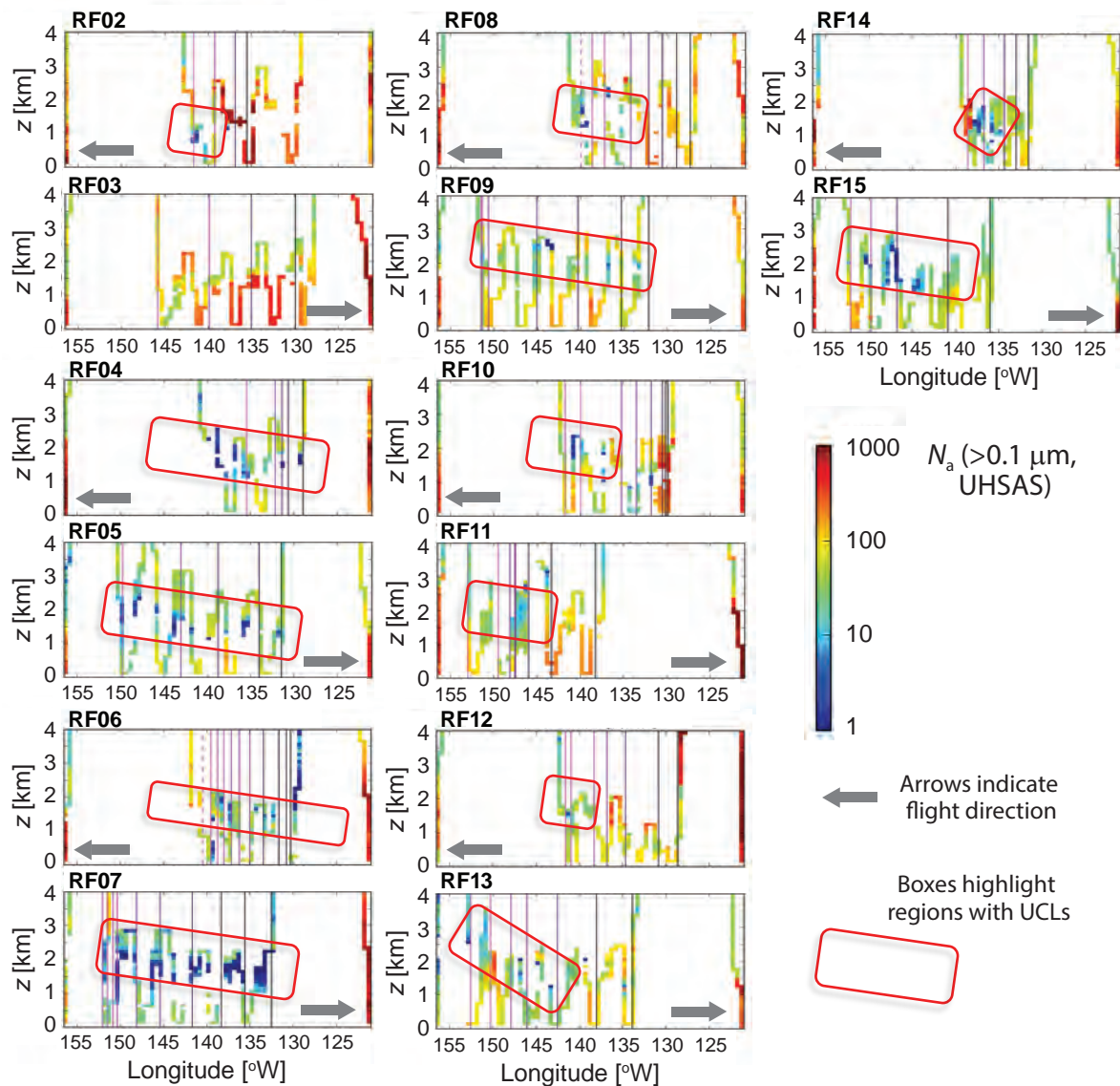
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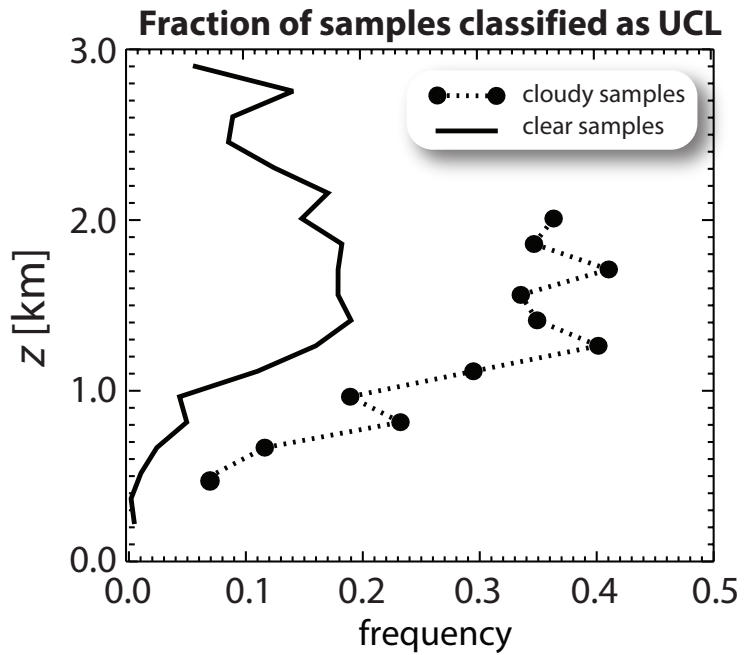
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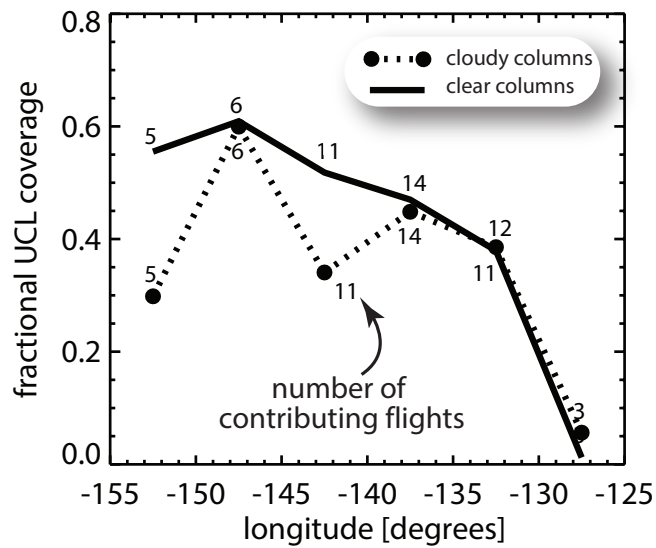
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1086 base height for optically thin clouds are 625 m and 1700 m respectively, with a median of  
1087 1150 m. The red dashed line shows the cumulative frequency of the height at which *in situ*  
1088 UCL cloud samples below 3 km altitude are found. It was not possible to determine the cloud  
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1091 be  $\sim 100$  m lower than the height of the sample on average. The 10th and 90th percentiles  
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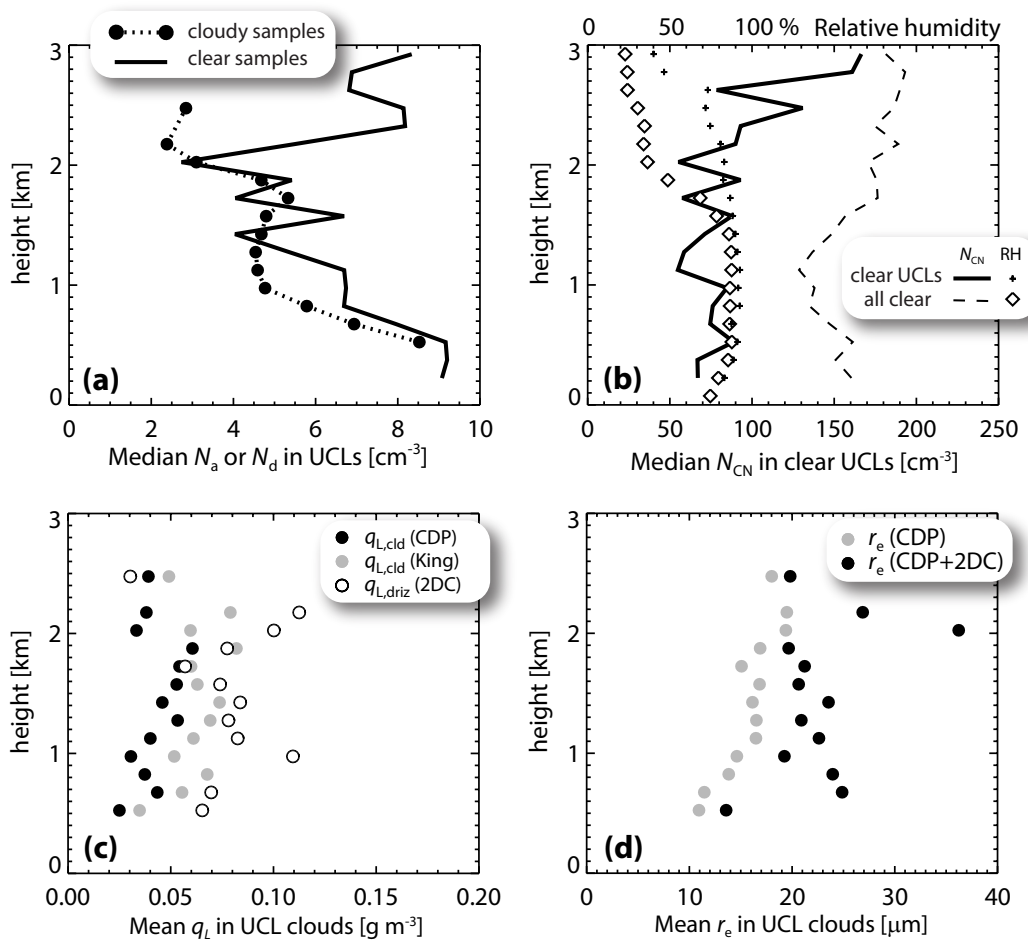
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 1100 return flights.



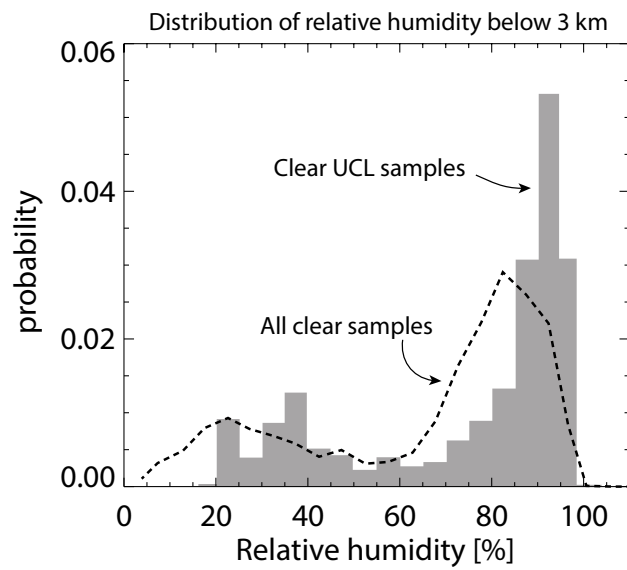
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 1102 down by whether the samples are clear or cloudy. Bins of 150 m altitude are used, and only those height levels  
 1103 containing at least thirty 1 Hz samples on each of three or more flights are shown.



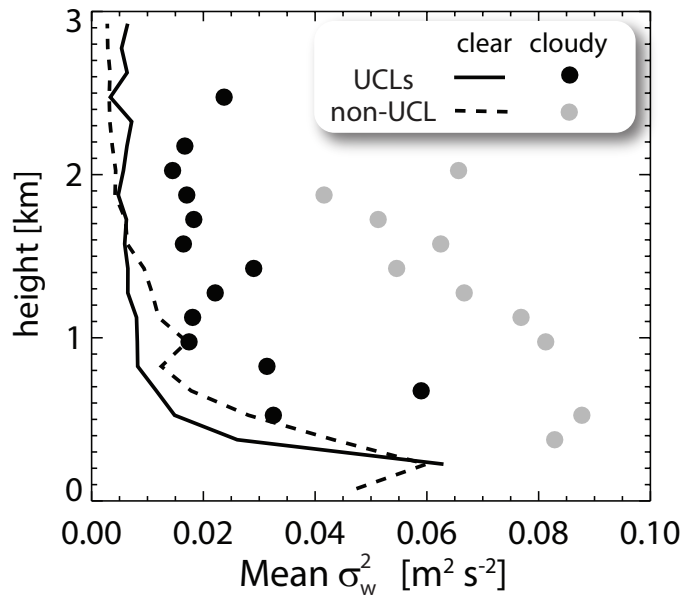
1104 FIG. 3. Conditional probability (given cloudy or clear column) that column contains a UCL at some level  
 1105 (below 3 km) assuming maximum overlap. The number of flights contributing to each longitude bin's statistics  
 1106 is shown next to the data values.



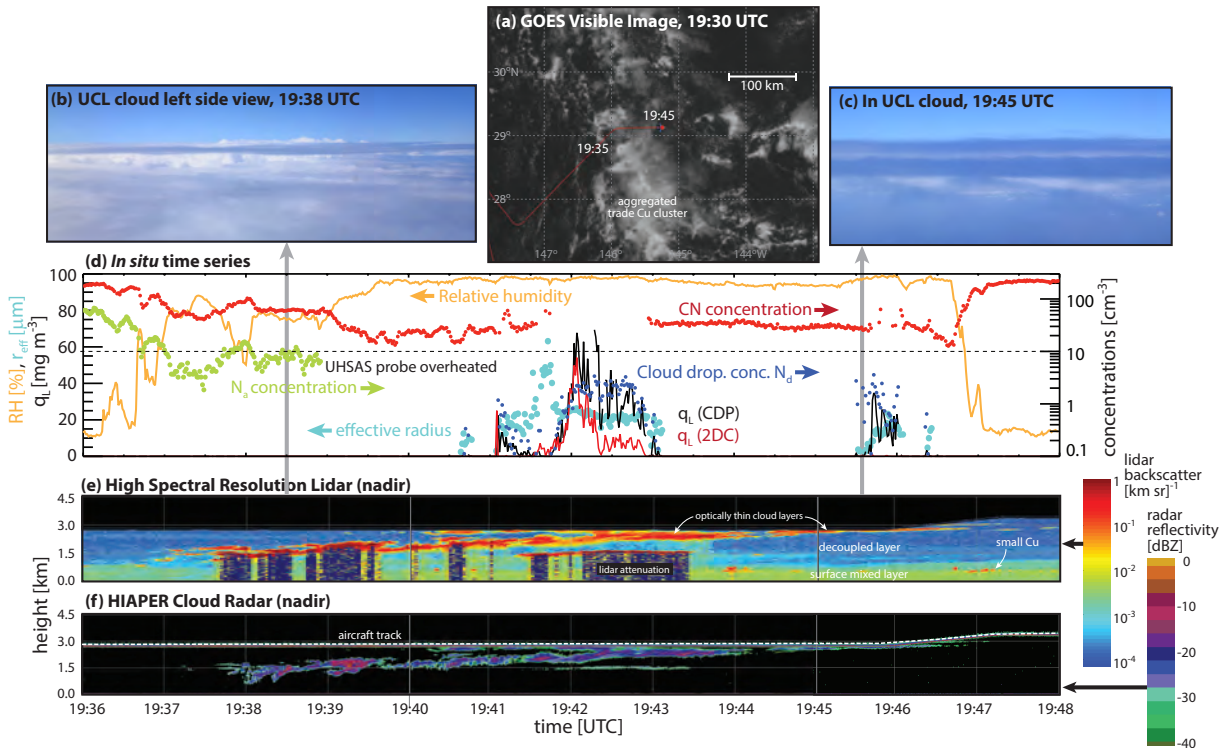
1107 FIG. 4. Composite height-resolved properties of UCLs from G-V *in situ* measurements during CSET. (a)  
 1108 Median concentrations of aerosol particles larger than  $0.1 \mu\text{m}$  ( $N_a$ ) in clear UCLs (solid) and of cloud droplets  
 1109 ( $N_d$ ) in UCL clouds, plotted as a function of UCL height; (b) Median condensation nucleus concentration (all  
 1110 particles larger than  $10 \text{ nm}$ ) in clear UCLs (solid) and in all clear samples (dashed). Also shown is the mean  
 1111 relative humidity for UCLs (small crosses) and for all clear samples (diamonds); (c) Mean liquid water contents  
 1112 measured with different probes (black solid circles: CDP probe, droplets with radii smaller than  $25 \mu\text{m}$ ; gray  
 1113 solid circles: King hotwire probe; open circles: 2DC shadowprobe, measuring larger droplets with radii larger  
 1114 than  $37.5 \mu\text{m}$ ); (d) Mean cloud drop effective radius ( $r_e$ ) considering only droplets with radii  $<25 \mu\text{m}$  (gray  
 1115 circles) and including all droplets (black circles).



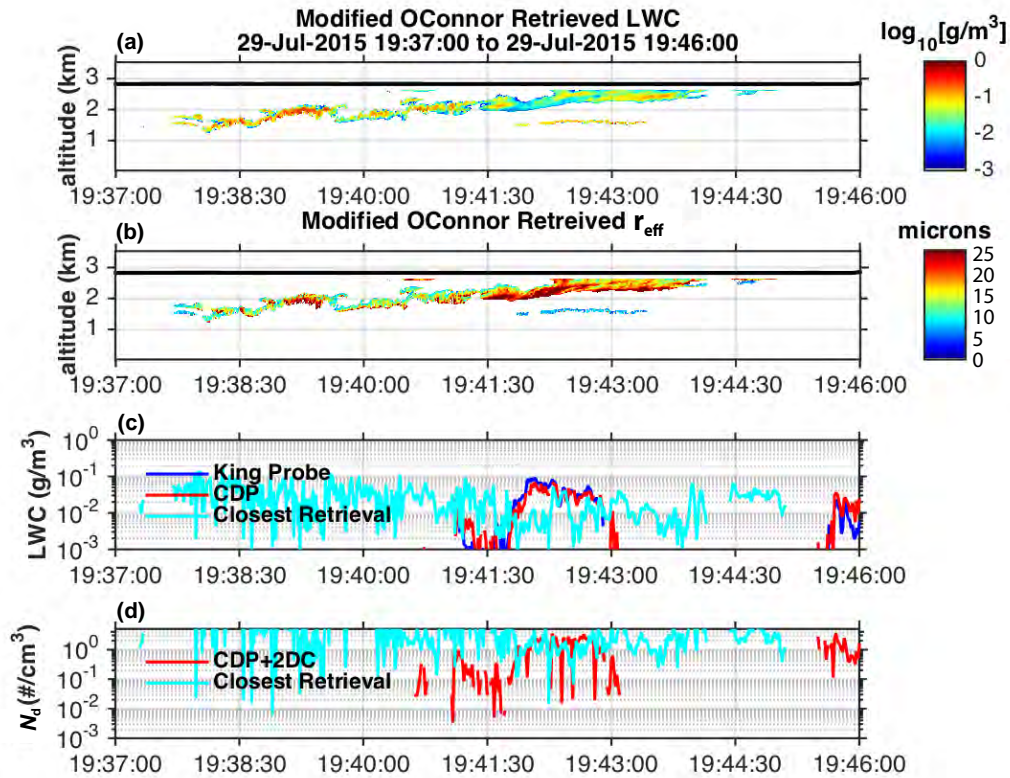
1116 FIG. 5. Probability distribution functions of clear sky relative humidity for all 1 Hz samples below 3 km  
 1117 (dashed line) and for all 1 Hz samples classified as clear UCLs (gray shading).



1118 FIG. 6. Height resolved composite mean values of 1 Hz turbulent vertical wind variance for both cloudy  
 1119 (circles, black=UCL; gray=non UCL) and clear (solid=UCL; dashed=non UCL) samples from CSET.

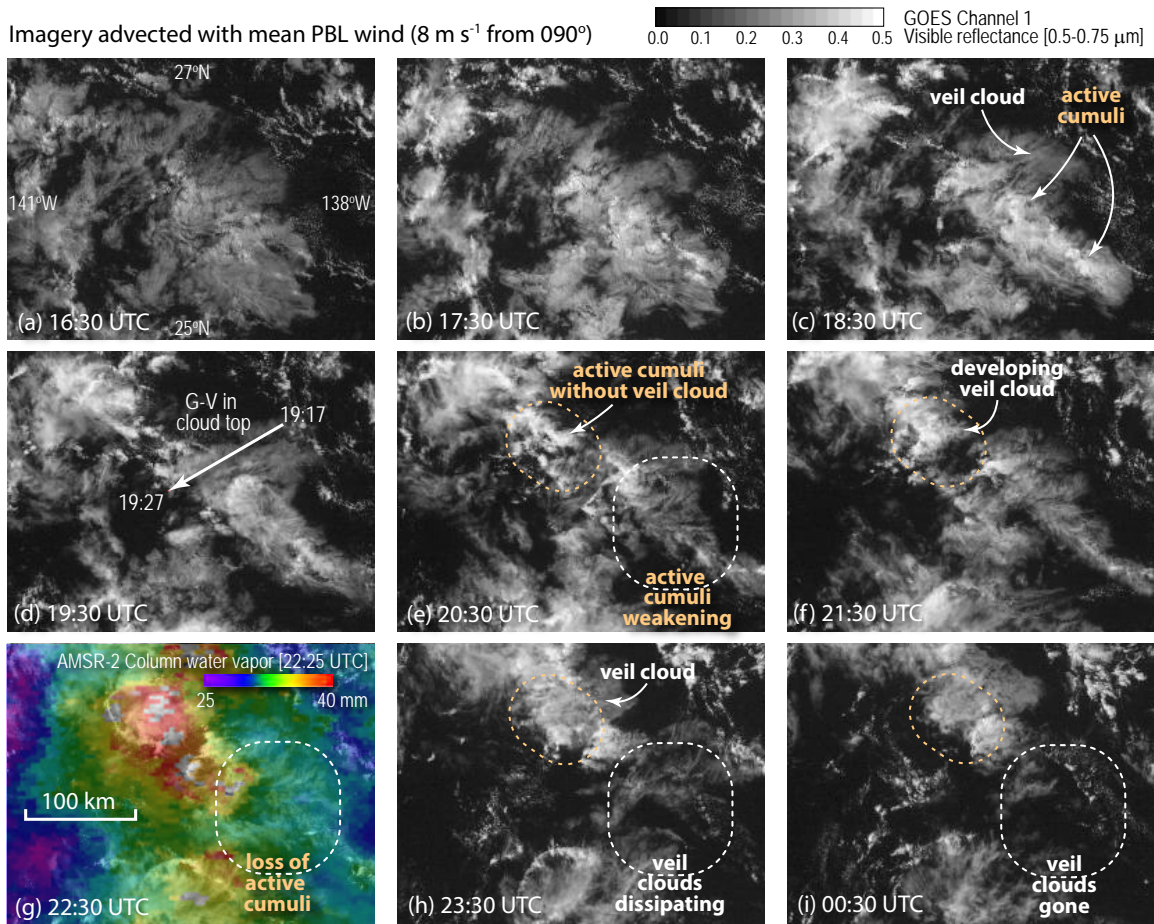


1120 FIG. 7. Case study of UCL cloud and clear layers sampled by the G-V between 19:36 and 19:48 during flight  
 1121 RF11 on 29 July 2015. (a) Visible satellite image from GOES-15 at 19:30 UTC showing aggregated Cu cluster  
 1122 approximately 100 km across along with the aircraft flight track (times marked on track); (b) photograph from  
 1123 the left side of the aircraft at 19:38 UTC showing bright Cu clouds and extensive layer (veil) cloud with gray  
 1124 appearance; (c) photograph taken during passage through optically thin UCL cloud at 19:45 UTC showing two-  
 1125 layered cloud structure at the level of the aircraft; (d) time series of several variables obtained using *in situ* G-V  
 1126 measurements during passage through system at 2.75 km altitude. Shown are concentrations of aerosol particles  
 1127 larger than  $0.1 \mu\text{m}$  ( $N_a$ ) (green, only available until 19:39 due to UHSAS overheating); CN concentration (red  
 1128 dots);  $N_d$  (blue dots); relative humidity (orange); liquid water content from the CDP (black line) and from the  
 1129 2DC (red line); effective radius (cyan); (e) lidar backscatter curtain below the aircraft from the high spectral  
 1130 resolution lidar (HSRL); (f) radar reflectivity factor curtain from the HIAPER cloud radar (HCR).

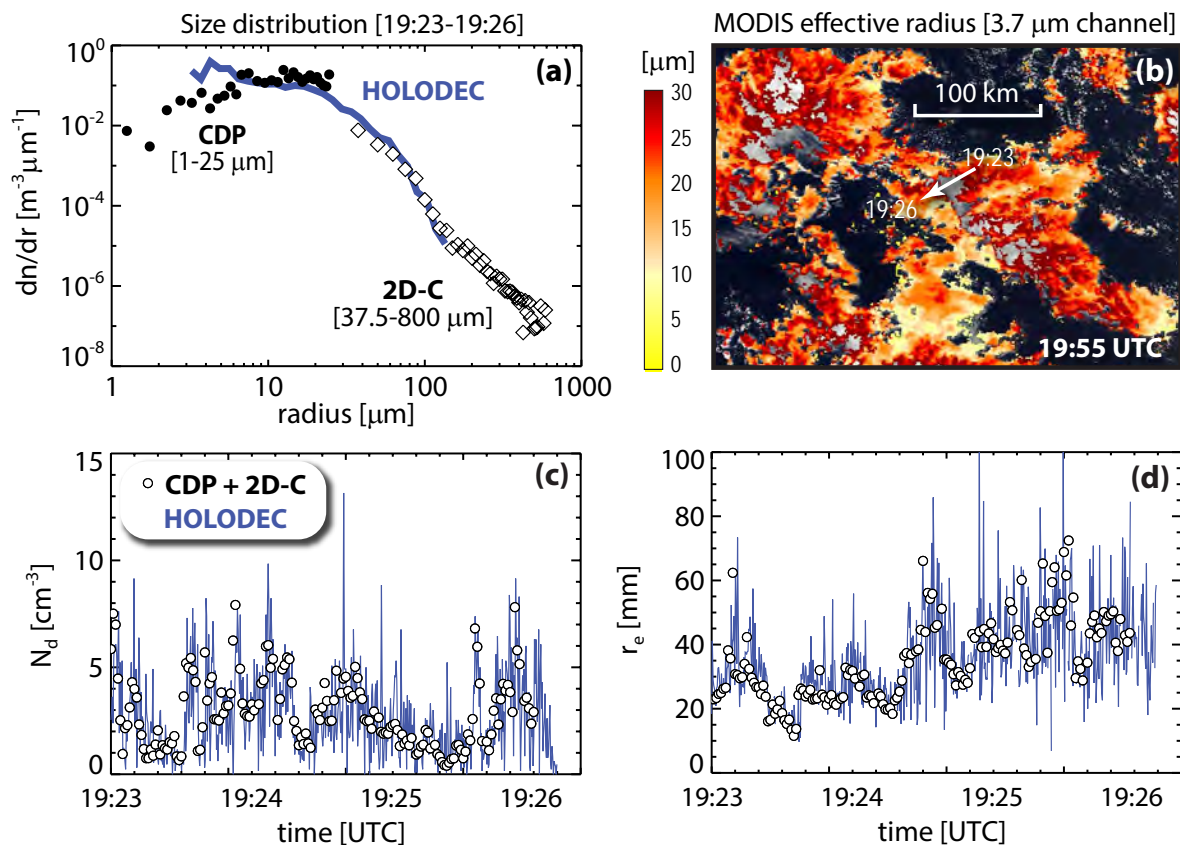


1131 FIG. 8. Radar-lidar retrievals of veil cloud properties indicate a low number concentration of large drops.  
 1132 This case is the same transect shown in Fig. 7. Here, we show liquid water content and effective radius (upper  
 1133 two panels) retrieved using a modified form (Schwartz and Ghate, 2017) of the O'Connor et al. (2005) method  
 1134 that was designed initially to quantify drizzle properties falling below stratocumulus cloud bases. The aircraft  
 1135 flight altitude is shown with the thick solid line. The lower two panels show comparisons of *in situ* liquid water  
 1136 content and droplet number concentration  $N_d$  with the closest retrieval possible ( $\sim 100$  m below the aircraft).

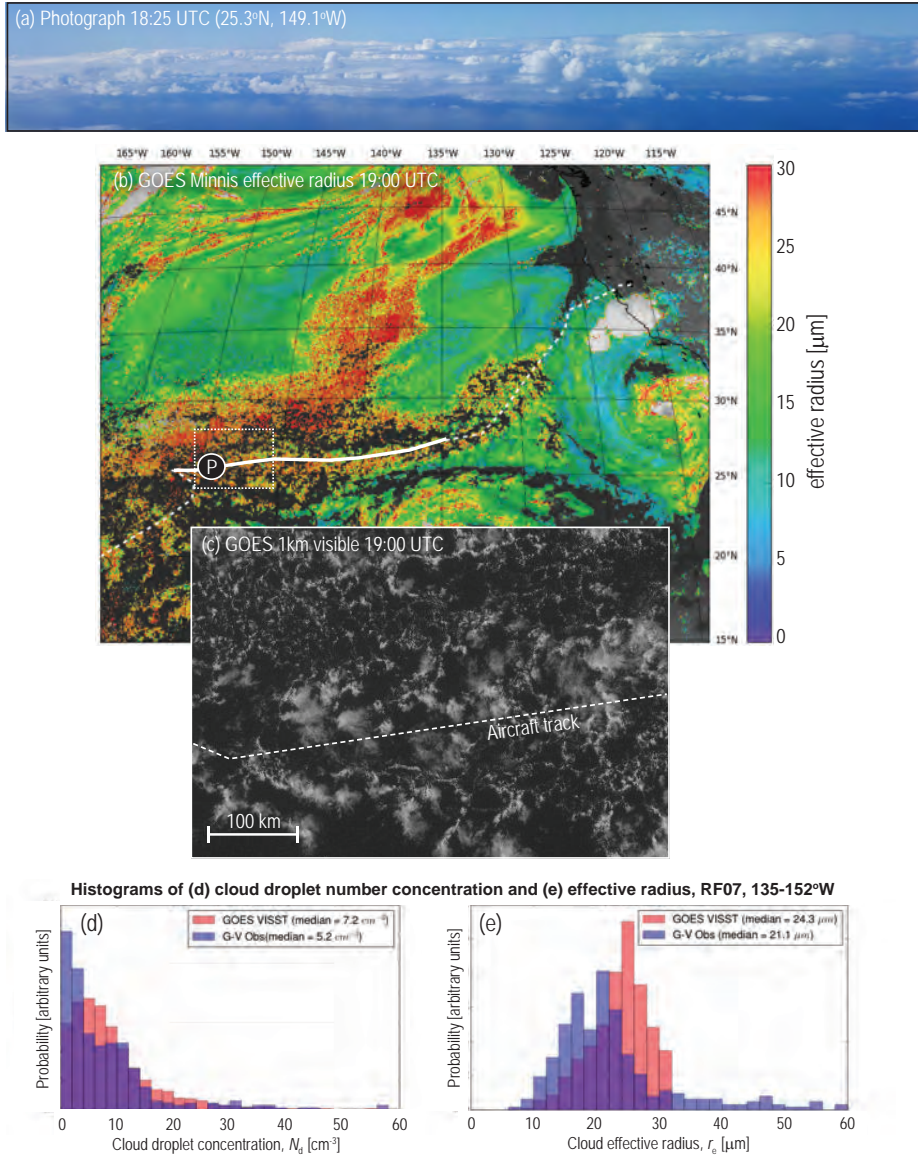
## GOES 1 km visible imagery for 27 July 2015 (CSET RF10)



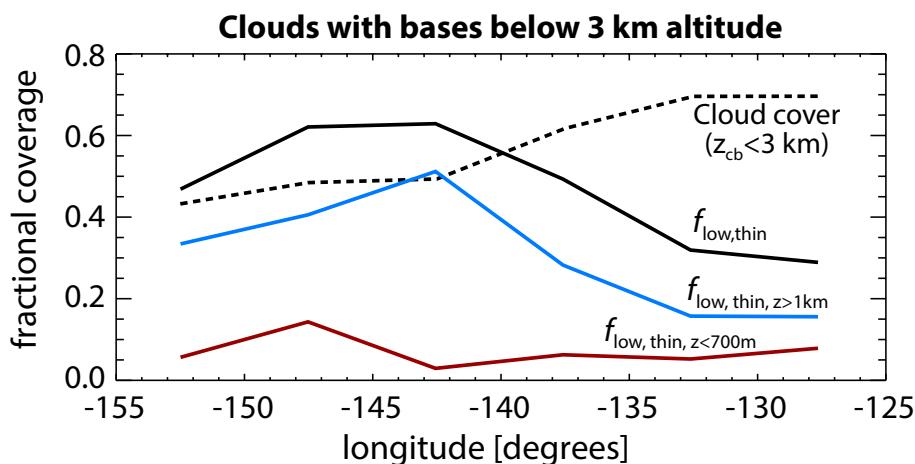
1137 FIG. 9. Evolution of active cumuli and veil clouds on 27 July (CSET RF10 flight) seen in GOES visible  
 1138 imagery over a period of eight hours. GOES visible imagery at  $\sim 1 \text{ km}$  resolution over a domain of approximately  
 1139  $3^\circ$  longitude by  $2^\circ$  latitude that is advected westward with the G-V observed mean PBL wind. The G-V flew a  
 1140 straight leg through relatively optically thin veil clouds at 2250-2300 m altitude from 19:17-19:27 UTC (panel d),  
 1141 where cloud droplet concentration varied from 30-40  $\text{cm}^{-3}$  during the first third of the leg (19:17-19:20 UTC)  
 1142 before dropping to values in the UCL range 1-10  $\text{cm}^{-3}$  after 19:22 UTC (Fig. 10). The time-lapse imagery  
 1143 demonstrates that optically thin stratiform veil clouds form in regions adjacent to clusters of optically thick  
 1144 (active) cumulus clouds. Initially, bright Cu form without surrounding stratiform clouds, but veil clouds tend  
 1145 to develop 1-2 hours later (panel e and f). When the driving cumulus weaken and dissipate, as occurs after  
 1146 20:30 UTC (panel e), the veil clouds persist for some hours before themselves dissipating (panel h and i). Panel  
 1147 g shows an overlay of the column water vapor from the AMSR-2 passive microwave imager on the GCOM-W1  
 1148 satellite that flies in the A-Train. The GCOM-W1 overpass time of 22:25 UTC is almost coincident with the  
 1149 22:30 GOES image and shows the aggregated Cu clusters tend to be associated with quite large moisture  
 1150 perturbations.



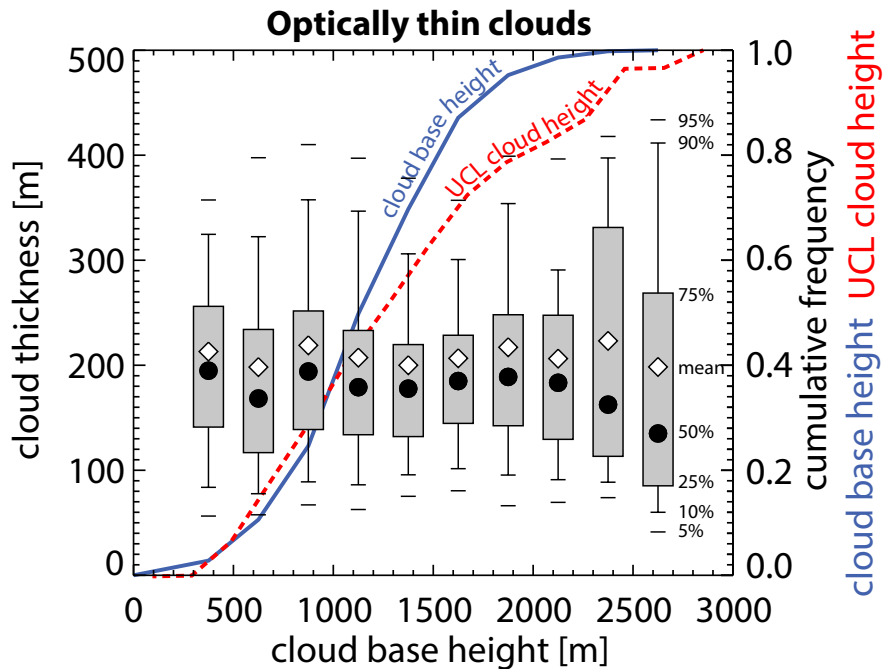
1151 FIG. 10. In situ and remotely sensed cloud microphysical properties near the top of the 27 July aggregated  
 1152 cumulus system shown in Fig. 9. *In situ* data shown here are from a section of the straight leg between 19:23  
 1153 and 19:26 UTC (a horizontal distance of about 30 km). (a) a broad mean size distribution was observed by  
 1154 the cloud probes on the G-V, with the HOLODEC probe size distribution showing excellent agreement. The  
 1155 effective radius of this size distribution is 37  $\mu m$ ; (b) retrieved effective radius map (3.7  $\mu m$  channel) from  
 1156 the MODIS instrument on the NASA Terra satellite showing prevalence of effective radii exceeding 25  $\mu m$ .  
 1157 Note that retrievals of cloudy pixels with effective radii larger than 30  $\mu m$  frequently fail (see section 5c) and  
 1158 constitute many of the gray regions; time series of (c) cloud droplet concentration  $N_d$  and (d) effective radius  
 1159 derived from the *in situ* CDP and 2D-C (circles) and from HOLODEC (blue line) showing excellent agreement.  
 1160 Effective radii as high as 40-60  $\mu m$  are found in this system.



1161 FIG. 11. Example of extensive patches of thin clouds associated with small cumulus clusters sampled by the  
 1162 G-V for over 1500 km (17:50-21:20 UTC) on flight RF07 on 19 July 2015. (a) a photograph from the left side of  
 1163 the aircraft at 18:25 UTC shows a cluster of bright Cu clouds interspersed with horizontally-extensive stratiform  
 1164 *veil* clouds; (b) satellite estimated cloud effective radius  $r_e$  from GOES-15 at 19:00 UTC (the white symbol on a  
 1165 circular black background shows the location of the photograph); (c) a visible GOES image at higher resolution  
 1166 of the region denoted by the box on the effective radius map showing trade Cu clusters in the form of loosely  
 1167 organized mesoscale open cells; the lower two panels show a statistical comparison of histograms of  $N_d$  (panel  
 1168 d) and  $r_e$  (panel e) from GOES-15 and from the G-V *in situ* data taken from the 1500 km portion of the G-V  
 1169 flight indicated by the thick solid line in (b) where the G-V was conducting low level sampling. See text for  
 1170 details of the data used to construct the histograms.



1171 FIG. 12. Fraction of columns containing optically thin low clouds increases westward. Figure shows estimates  
 1172 from the HSRL in zenith-pointing mode on the G-V aircraft from near-surface level legs using the cloud mask of  
 1173 Schwartz and Ghate (2017). Only clouds with bases below 3 km are included. Dashed line: low cloud fractional  
 1174 cover from all clouds detected by the HSRL. Solid black line: fraction of low clouds that are optically thin,  
 1175 i.e. clouds detected by the HSRL that do not fully attenuate the lidar (see Leahy et al., 2012). Optically thin  
 1176 low clouds are further divided into clouds with bases above 1 km altitude (blue line) and those near the lifting  
 1177 condensation level ( $z < 700$  m, red solid line).



1178 FIG. 13. Geometric thickness of optically thin low clouds sampled by the HSRL on the G-V aircraft during  
 1179 CSET, plotted against cloud base height (box-whiskers). Optically thin cloud layers are those that do not fully  
 1180 attenuate the HSRL beam, which is assessed using the HSRL signal to noise ratio. Here, we only include data  
 1181 from subcloud straight and level legs where the HSRL was nadir viewing. Box-whisker plots show percentiles  
 1182 of cloud thickness (5th, 10th, 25th, 75th, 90th, 95th percent) using box-whiskers, with black circles indicating  
 1183 medians and diamonds means. The solid blue line shows the cumulative frequency of cloud base height for  
 1184 optically thin clouds, indicating that the 10th and 90th percentiles of cloud base height for optically thin clouds  
 1185 are 625 m and 1700 m respectively, with a median of 1150 m. The red dashed line shows the cumulative  
 1186 frequency of the height at which *in situ* UCL cloud samples below 3 km altitude are found. It was not possible to  
 1187 determine the cloud base height of the UCL *in situ* cloud samples, but their low geometrical thickness suggests  
 1188 that the cloud base height associated with any given UCL *in situ* cloud sample would only be  $\sim 100$  m lower  
 1189 than the height of the sample on average. The 10th and 90th percentiles of UCL cloud height are 560 m and  
 1190 2300 m respectively, with a median of 1250 m.