

**1 Observational evidence for a negative shortwave**  
**2 cloud feedback in mid to high latitudes**

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**Key Points.**

- A negative shortwave cloud feedback is observed in mid to high Southern latitudes.
- This negative feedback results from increasing cloud optical depth with temperature.
- Models are in qualitative agreement with observations in mid to high latitudes.

**Abstract.**

Exploiting the observed robust relationships between temperature and optical depth in extratropical clouds, we calculate the shortwave cloud feedback from historical data, by regressing observed and modeled cloud property histograms onto local temperature in mid to high Southern latitudes. In this region, all CMIP5 models and observational data sets predict a negative cloud feedback, mainly driven by optical thickening. Between 45° and 60° S, the mean observed shortwave feedback ( $-0.91 \pm 0.82 \text{ W m}^{-2} \text{ K}^{-1}$ , relative to local rather than global-mean warming) is very close to the multi-model mean feedback in RCP8.5 ( $-0.98 \text{ W m}^{-2} \text{ K}^{-1}$ ), despite differences in the meridional structure. In models, historical temperature–cloud property relationships reliably predict the forced RCP8.5 response. Because simple theory predicts this optical thickening with warming, and cloud amount changes are relatively small, we conclude that the shortwave cloud feedback is very likely negative in the real world at mid to high latitudes.

## 1. Introduction

18 The cloud feedback has been identified as the dominant source of uncertainty in model-  
19 based estimates of climate sensitivity, primarily because of the shortwave radiation re-  
20 sponse associated with tropical low clouds [Boucher *et al.*, 2013]. To a large extent,  
21 the uncertain cloud-radiative response reflects difficulties in representing the effects of  
22 small-scale processes on the cloud water budget in coarse climate model grids using pa-  
23 rameterizations. Among the most uncertain parameterized processes are those related  
24 to convective mixing [Zhao, 2014; Sherwood *et al.*, 2014; Webb *et al.*, 2015] as well as  
25 ice-phase cloud microphysics [Storelvmo *et al.*, 2015, and references therein].

26 Despite these large uncertainties persisting across generations of climate models, some  
27 robust signals emerge. In terms of the shortwave (SW) cloud feedback, current models  
28 agree on a negative feedback in mid to high latitudes, mainly caused by an optical thicken-  
29 ing and brightening of the clouds [Zelinka *et al.*, 2012a; Gordon and Klein, 2014; Ceppi  
30 *et al.*, 2016]. Warming-induced phase changes in mixed-phase clouds (dominant in mid  
31 to high latitudes) are believed to be an important driver of this optical thickening, at  
32 least in models [Tsushima *et al.*, 2006; McCoy *et al.*, 2014; Ceppi *et al.*, 2016], although  
33 increases in the “adiabatic” cloud water content could also contribute [Somerville and  
34 Remer, 1984; Betts and Harshvardhan, 1987; Tselioudis *et al.*, 1992]. In models, the re-  
35 lationship between optical depth and temperature remains similar across timescales, so  
36 that the forced optical depth response in global warming experiments is well-predicted by  
37 unforced seasonal or interannual fluctuations [Gordon and Klein, 2014]. This supports  
38 the idea that the cloud optical depth increase in high latitudes is a direct response to

39 warming, and suggests that the associated negative SW feedback might be predictable  
40 from historical data. To our knowledge, however, the robust optical depth–temperature  
41 relationships have not been exploited thus far to predict the SW cloud feedback in models  
42 and observations.

43 The purpose of this paper is to demonstrate that the cloud water increase, and the  
44 associated optical thickening and negative SW feedback in mid to high latitudes, can  
45 all be detected from historical data in both models and observations in the Southern  
46 Hemisphere. Furthermore, the SW cloud feedback in the RCP8.5 experiment is well-  
47 predicted from historical model simulations in mixed-phase regions, consistent with the  
48 time scale-invariance of optical depth–temperature relationships found in previous studies.  
49 While observational uncertainties and disagreements between satellite products limit our  
50 ability to produce accurate quantitative estimates of the SW cloud feedback in the real  
51 world, we will show that models and observations are in qualitative agreement on this  
52 negative SW cloud feedback.

## 2. Data and Methods

53 This study combines observed and modeled cloud property data, all in monthly-mean  
54 resolution. We first briefly describe the satellite observations. Liquid water path changes  
55 are assessed using 20 years of satellite microwave retrievals [UWisc data set; *O’Dell et al.*,  
56 2008], covering the period January 1989 to December 2008. Additionally, we use cloud  
57 amount retrievals, available from three satellite-based data sets: ISCCP [*Rossow and*  
58 *Schiffer*, 1999], MISR [*Diner et al.*, 1998] and MODIS [*King et al.*, 2003; *Platnick et al.*,  
59 2003], providing 25, 12, and 14 years of data, respectively. These cloud amount retrievals  
60 have been binned into cloud top pressure (or height) versus optical depth histograms.

61 For details on the preparation of these simulator-oriented data sets, see *Marchand et al.*  
62 [2010], *Zhang et al.* [2012], *Pincus et al.* [2012], and *Marchand* [2013], as well as Text S1.

63 The model data used here include output from the historical and RCP8.5 experiments,  
64 where the RCP8.5 response or feedback is based on differences between 1981–2000 and  
65 2081–2100. The 30 models included in this study are listed in the Supporting Information,  
66 Table S1. Only a subset of these models provide cloud property histograms for comparison  
67 with observations, and due to limited availability of historical cloud histogram data we  
68 also use AMIP output (see Table S1). These cloud histograms are produced using satellite  
69 simulators, which mimic the cloud properties that would be retrieved by a satellite if the  
70 modeled clouds existed in the real world; for details, see *Klein and Jakob* [1999], *Webb*  
71 *et al.* [2001], and *Klein et al.* [2013]. We have verified correct simulator implementation  
72 in these models as in *Zelinka et al.* [2012b] (see Text S1 for details).

73 As outlined in the Introduction, the main goal of this study is to demonstrate the  
74 existence of a negative cloud optical depth feedback in mid to high latitudes that is  
75 detectable in the context of unforced seasonal or interannual variability. Assuming this  
76 feedback is mainly driven by the direct effect of local warming, we estimate the SW cloud  
77 feedback in two steps. In the first step, we regress cloud property histograms onto local  
78 lower-tropospheric temperature (defined as the 500–850 hPa layer mean) in models and  
79 observations. The choice of this pressure range is based on the fact that the bulk of cloud  
80 water is typically contained in this layer in models [see e.g. *Komurcu et al.*, 2014; *Ceppi*  
81 *et al.*, 2016]. For models, we use 1981–2000 historical or AMIP data; for observations, the  
82 satellite data are regressed onto ERA-Interim reanalysis temperature using the full length  
83 of each of the satellite products. The regressions are calculated at each latitude, using

84 data for all months and longitudes linearly interpolated onto the cloud histogram grid.  
85 Prior to the regression analysis, we remove the annual-mean value at every gridpoint, and  
86 average the data over non-overlapping, 20°-wide longitude boxes. This last step ensures  
87 that the cloud anomalies are nearly uncorrelated between adjacent longitude points, and  
88 can be treated as independent realizations in both time and longitude space, which is  
89 necessary for an accurate estimation of confidence intervals for the regression slopes.

90 Note that since some of the satellite instruments do not report values over land or sea ice  
91 gridpoints, only ocean gridpoints are included in the regression analysis, and we restrict  
92 the analysis to the Southern Hemisphere where most midlatitude areas are ocean-covered.  
93 Furthermore, because instruments measuring solar reflectance (such as ISCCP, MISR, and  
94 MODIS) are known to produce large positive biases in cloud optical depth at high solar  
95 zenith angles [Loeb and Davies, 1996], for MISR and MODIS we exclude points with solar  
96 zenith angle  $> 60^\circ$  at the time of satellite overpass. For ISCCP, the results exhibit very  
97 little sensitivity to the exclusion of high solar zenith angle retrievals (not shown), so such  
98 retrievals are included in the analysis. The sensitivity of the results to these choices is  
99 discussed in the Supporting Information.

100 In the second step, the cloud property histogram regressions are converted to anomalous  
101 top-of-atmosphere SW radiative fluxes by multiplying with SW cloud-radiative kernels  
102 [described in Zelinka *et al.*, 2012b] and integrating over all 49 ( $7 \times 7$ ) cloud top pressure–  
103 optical depth bins. This yields a cloud feedback in units of  $\text{W m}^{-2} \text{K}^{-1}$ , which we call  
104 the “predicted” cloud feedback, and we compare it with the “actual” feedback in RCP8.5  
105 obtained by the approximate partial radiative perturbation method [APRP; Taylor *et al.*,  
106 2007]. Since the cloud-radiative kernels are functions of surface albedo, we use a 15-year

107 climatology (March 2000–February 2015) of CERES EBAF clear-sky surface upward and  
108 downward SW fluxes [Loeb *et al.*, 2009] to calculate the observed surface albedo. Also,  
109 because we assume that the cloud property–temperature relationships are independent of  
110 the month of the year, the cloud-radiative kernels are averaged over all calendar months  
111 prior to multiplication with the cloud histogram regression matrices.

### 3. Results

#### 3.1. Observed and modeled changes in LWP and cloud amount

112 We begin by assessing how temperature affects two key cloud properties relevant to  
113 SW radiation, cloud liquid water path (LWP) and cloud amount (or fractional coverage),  
114 in the mid to high Southern latitudes. Models and observations agree on a positive re-  
115 lationship between LWP and low-level temperature (Fig. 1a), although the magnitude  
116 of the relationship varies considerably among models. Compared to most models, the  
117 observed LWP–temperature relationship is weaker in magnitude poleward of about 47°  
118 S and stronger equatorward thereof, but is very highly statistically significant, and re-  
119 mains the same whether the seasonal cycle is removed or not (thin and thick red curves  
120 in Fig. 1a). Because cloud optical thickness is approximately linearly proportional to the  
121 LWP [Stephens, 1978], the positive LWP–temperature relationships imply optical thick-  
122 ening with warming. Assuming these relationships hold for the forced global warming  
123 case, and all other things remaining equal, one would thus expect brighter clouds and  
124 therefore a negative SW cloud feedback to occur in mid to high latitudes.

125 Note that although Fig. 1a–b shows gridbox-mean rather than in-cloud LWP values, the  
126 LWP increases are *not* due to cloud amount increases; the results remain qualitatively  
127 unchanged if the LWP values are normalized by cloud fraction before calculating the

128 response (not shown). Furthermore, the cloud amount and LWP changes are essentially  
129 uncorrelated across models over the 45°–60° S region ( $r = 0.06$ ). While cloud reflectivity  
130 is also affected by changes in ice water path (IWP, not shown here), the cloud ice response  
131 is substantially smaller than the LWP change in RCP8.5 [see Fig. 1 in *Ceppi et al.*, 2016],  
132 suggesting this is a second-order effect in models.

133 Another potential effect of clouds on SW radiation comes from cloud amount changes.  
134 While the cloud amount sensitivity to low-level temperature is very model-dependent, the  
135 mean model behavior is to increase cloud cover poleward of about 50° S with warming,  
136 with a weak decrease equatorward thereof (Fig. 1c). By contrast, all three satellite instru-  
137 ments show a weak but statistically significant cloud amount increase with warming at all  
138 latitudes poleward of about 40° S. Although they agree on the sign of the response, quan-  
139 titative differences exist, with MODIS systematically indicating the largest cloud fraction  
140 increases. It should be noted that in observations, low-level temperature is highly corre-  
141 lated with lower-tropospheric stability [as measured by the estimated inversion strength;  
142 *Wood and Bretherton*, 2006] over the Southern midlatitudes (not shown). Since low-level  
143 stability is an important control on low cloud amount [*Klein and Hartmann*, 1993; *Wood*  
144 *and Bretherton*, 2006], the observed positive cloud amount–temperature relationship may  
145 in fact reflect the effect of boundary-layer stability on low cloud amount, rather than a  
146 direct effect of temperature. The influence of low-level stability on low cloud amount  
147 appears to be underestimated by models [*Qu et al.*, 2015].

148 The historical relationships between cloud properties and temperature are useful indica-  
149 tors of the cloud feedback only to the extent that they accurately predict future changes.  
150 So, are the predicted and actual cloud responses similar? The bottom row of Fig. 1 shows

151 the actual LWP and cloud amount response in the RCP8.5 experiment, for compari-  
152 son with the predicted response. Note that although the RCP8.5 response is not purely  
153 temperature-driven and also contains a direct CO<sub>2</sub> effect [*Sherwood et al.*, 2015], the cloud  
154 response in AMIP4K is very similar (not shown), suggesting that the response is mainly  
155 warming-induced. Overall, the actual responses are remarkably similar to those predicted  
156 from historical model data in an ensemble-mean sense. Comparing the responses across  
157 models, we find that the predicted and actual LWP changes are well-correlated over the  
158 45°–60° S latitude range and close to the one-to-one line (Fig. 2a), suggesting that future  
159 LWP changes are reasonably well-predicted by historical relationships. In this and follow-  
160 ing scatterplots, we use the uncertainty in the relationship between predicted and actual  
161 response in models to derive observational confidence intervals, such that the confidence  
162 interval width is proportional to the standard deviation of the residuals relative to the  
163 one-to-one line (Text S3).

164 The relationship between actual and predicted cloud amount change is also positive,  
165 but the agreement is weaker than for LWP (Fig. 2b). However, we will show that  
166 increasing optical depth, rather than cloud amount, is the main driver of the negative  
167 SW cloud feedback simulated by models in mid to high latitudes. The scatterplots in  
168 Fig. 2 demonstrate that historical, seasonal relationships between cloud properties and  
169 local temperature are representative of the forced, long-term cloud response to future  
170 warming. They also illustrate that in the 45°–60° S region, models generally overestimate  
171 the LWP increase compared to observations (vertical red bar in Fig. 2a), in some cases  
172 by a considerable amount. By contrast, the predicted cloud amount change tends to be

173 less positive than in observations, although the difference is small. The implications of  
174 these differences on the SW cloud feedback will be discussed in the next subsection.

### 3.2. Observed and modeled SW cloud feedback

175 To estimate the SW cloud feedback in models and observations, under the assumption  
176 the feedback is mainly driven by local temperature changes, we proceed as in the previous  
177 section and regress the cloud amount histograms onto low-level temperature. Given the  
178 robust increases in LWP seen in the mid to high Southern latitudes, we expect to find a  
179 shift in the cloud amount histogram toward higher optical depth as temperature increases.  
180 As illustrated in Fig. 3, this is indeed the case: over the  $45^{\circ}$ – $60^{\circ}$  S region, the cloud amount  
181 response mainly consists of a dipole along the optical depth dimension, reflecting a shift  
182 of the climatological cloud distribution toward higher  $\tau$  values. The cloud histogram  
183 responses in MISR (Fig. S1), MODIS (Fig. S2), and the model ISCCP simulators (not  
184 shown) are all qualitatively similar. The cloud histogram regression matrices (illustrated  
185 in Fig. 3b) are multiplied with the SW cloud-radiative kernel at each latitude to yield a  
186 predicted SW cloud feedback (see Methods).

187 As expected from the results above, models and observations predict a negative SW  
188 cloud feedback poleward of about  $45^{\circ}$  S, coincident with the region of increasing LWP  
189 with temperature (Fig. 4a). While different observational data sets disagree on the mag-  
190 nitude of the negative feedback, especially at the highest latitudes, they agree on the sign  
191 and overall latitudinal structure. They are also consistent with the bulk of the model dis-  
192 tribution, although the observed negative feedback pattern appears to be shifted toward  
193 lower latitudes compared with most models. This shift agrees qualitatively with the differ-  
194 ences in the LWP and cloud amount responses to warming (cf. Fig. 1a–b). The predicted

195 cloud feedback is remarkably similar to the actual RCP8.5 cloud feedback in this set of  
196 climate models both in terms of magnitude and meridional structure (Fig. 4b), and the  
197 values are well-correlated across models in the 45°–60° S latitude band (Fig. 5), confirm-  
198 ing the idea that historical temperature–cloud brightness relationships are representative  
199 of long-term changes [*Gordon and Klein, 2014*].

200 Averaging the three observational estimates together, we find a mean observed SW  
201 feedback of  $-0.91 \pm 0.82 \text{ W m}^{-2} \text{ K}^{-1}$ , significantly negative, and close to the mean ac-  
202 tual feedback in RCP8.5 ( $-0.98 \text{ W m}^{-2} \text{ K}^{-1}$ ). It should be noted, however, that the  
203 close agreement between observations and models in Fig. 5 masks disagreements in the  
204 meridional structure of the cloud feedback, as described in the previous paragraph. Such  
205 disagreements in the meridional structure of the SW cloud feedback may have important  
206 implications for the models’ ability to correctly simulate the spatial distribution of the  
207 temperature response.

208 Using the cloud property histograms allows us to decompose the cloud feedback into  
209 effects of cloud optical depth and cloud amount changes (Fig. S3), following the method  
210 of *Zelinka et al. [2013]*. In general, the optical depth increase explains most of the neg-  
211 ative SW cloud feedback in observations and for the multi-model mean, although cloud  
212 amount changes do contribute substantially to the inter-model spread in SW feedback.  
213 For MODIS, the optical depth and cloud amount effects appear to be of comparable  
214 magnitude poleward of 45° S, while the other two satellite data sets predict a larger  
215 (more negative) optical depth feedback. Hence, Fig. S3 shows that the disagreement in  
216 the magnitude of the observed negative cloud feedback in Southern midlatitudes between  
217 satellite products is mainly associated with the optical depth effect rather than with cloud

218 amount changes. Note that there are two additional terms in this decomposition of the  
219 cloud histogram response, one reflecting the effect of cloud altitude changes and another  
220 representing a residual term [see *Zelinka et al.*, 2013]; both are very small (not shown).

#### 4. Discussion

221 The results in the previous section have shown that in mid to high Southern latitudes,

222 1. cloud optical depth increases with warming are detectable in observations and his-  
223 torical model simulations,

224 2. in models the historical seasonal optical depth–temperature relationships are good  
225 predictors of the future SW cloud feedback, and

226 3. the predicted negative SW cloud feedback is qualitatively similar in models and  
227 observations.

228 The ubiquity of this negative feedback across models and observations is likely at least  
229 in part a result of robust phase changes in mixed-phase cloud regions, which cause in-  
230 creases in LWP and optical depth with warming [*Tsushima et al.*, 2006; *McCoy et al.*,  
231 2014; *Storelvmo et al.*, 2015; *Ceppi et al.*, 2016]. While the possible importance of the  
232 phase change effect in the real world is difficult to demonstrate due to limitations in the  
233 availability and quality of cloud phase observations, changes in microphysical phase con-  
234 version rates have been shown to be the main driver of the negative optical depth feedback  
235 in models [*Ceppi et al.*, 2016]. Differences in the parameterization of microphysical phase  
236 change processes also likely account for at least part of the large inter-model differences in  
237 LWP and cloud optical depth sensitivity to warming [*Komurcu et al.*, 2014; *McCoy et al.*,  
238 2015; *Cesana et al.*, 2015].

239 Although the three satellite data sets are in qualitative agreement on an optical thick-  
240 ening (and associated negative SW cloud feedback) in high Southern latitudes, and are in  
241 relatively close agreement in a regional mean sense (the 45°–60° S–mean feedback ranging  
242 between  $-0.76 \text{ W m}^{-2} \text{ K}^{-1}$  for MODIS and  $-1.07 \text{ W m}^{-2} \text{ K}^{-1}$  for MISR), the differences  
243 between data sets are not negligible locally (cf. Figs. 4 and S3), reflecting disagreements  
244 in the cloud optical depth response to warming (Fig. S4). We believe the differences  
245 between observational data sets result from relatively large uncertainties in the satellite  
246 retrievals of cloud properties, with the uncertainty sources being specific to each data set.  
247 Some known error sources, and their possible impacts on our results, are discussed in the  
248 Supporting Information (Text S2). It should be kept in mind that measurement errors  
249 due to illumination and viewing angle, for example, are not included in the instrument  
250 simulators in climate models. For this reason, instrument simulators may better reflect  
251 the clouds in the models than real satellite observations characterize clouds in nature.

252 In addition to the uncertainty associated with errors in satellite retrievals, further un-  
253 certainty in the magnitude of the real SW cloud feedback results from the imperfect pre-  
254 diction of future responses from historical model data, as illustrated in the scatterplots in  
255 Figs. 2 and 5; the calculation of the observational confidence intervals is derived from this  
256 uncertainty (Text S3). The differences between predicted and actual response result from  
257 effects not accounted for by the simple regression on local temperature; an obvious exam-  
258 ple would be the radiative effect of increasing  $\text{CO}_2$  concentrations, but other factors such  
259 as atmospheric circulation, lower-tropospheric stability, vertical and horizontal moisture  
260 fluxes, to name a few, are likely also contributing to the forced cloud response.

261 Despite the current shortcomings of cloud property observations, we believe that the  
262 positive optical depth–temperature relationships are real and physical for the following  
263 two reasons: (a) four independent observational data sets and all CMIP5 models agree  
264 on mixed-phase clouds becoming optically thicker (or equivalently, their water content  
265 increasing) with warming, and (b) in such cold clouds a positive optical depth–temperature  
266 relationship is *expected* from relatively basic physical temperature-related mechanisms  
267 (phase transitions, and increasing adiabatic water content). (Note that while optical  
268 depth increases linearly with LWP only assuming constant droplet radius [*Stephens*, 1978],  
269 a more realistic assumption of constant particle number also leads to higher optical depth  
270 as LWP increases.) With no indication of compensating large cloud amount decreases with  
271 warming in the real world, and strong observational and modeling evidence for optical  
272 depth increases, we conclude that the shortwave cloud feedback in a future warmer climate  
273 will very likely be negative in mid to high Southern latitudes.

## 5. Conclusions

274 Using historical CMIP5 model data and satellite retrievals of cloud properties, we have  
275 shown that as the atmosphere warms, cloud liquid water (and hence optical depth) con-  
276 sistently increase in mid to high latitudes (poleward of  $\sim 45^\circ$ ) in the Southern Hemi-  
277 sphere, with an additional weak cloud amount response to warming. Although models  
278 disagree on the magnitude of the cloud liquid water increase, it is present in all mod-  
279 els and in observations, and is supported by robust temperature-dependent mechanisms  
280 (phase changes in mixed-phase clouds, and adiabatic cloud water content increases). To  
281 estimate the SW radiation response associated with the optical thickening of the clouds,  
282 cloud property histograms binned by cloud top pressure and optical depth are regressed

283 on lower-tropospheric temperature and combined with cloud-radiative kernels [*Zelinka*  
284 *et al.*, 2012b]. Consistent with the cloud water response, all models and observational  
285 data sets predict a negative SW cloud feedback in mid to high Southern latitudes, and  
286 observations lie well within the model distribution. In the 45°–60° S latitude band, the  
287 predicted feedback in observations ranges between  $-0.76 \text{ W m}^{-2} \text{ K}^{-1}$  (MODIS) and  $-1.07$   
288  $\text{W m}^{-2} \text{ K}^{-1}$  (MISR), with an estimated mean value of  $-0.91 \pm 0.82 \text{ W m}^{-2} \text{ K}^{-1}$ , close to  
289 the actual feedback in RCP8.5 ( $-0.98 \text{ W m}^{-2} \text{ K}^{-1}$ ). For models, the feedback predicted  
290 from historical seasonal temperature variations is a good predictor of the actual feedback  
291 in the RCP8.5 experiment ( $r = 0.78$ ), supporting the idea that a warming-induced cloud  
292 optical depth increase is detectable in the historical record.

293 Observed cloud optical depth–temperature relationships in extratropical clouds have  
294 been proposed as a promising potential observational constraint on modeled cloud feed-  
295 backs by *Klein and Hall* [2015], who also note that such observational constraints must  
296 be supported by a robust, well-understood physical mechanism in order to be credible.  
297 In the context of the cloud optical depth feedback, an important question for future work  
298 is therefore to clarify the relative importance of phase change effects and adiabatic water  
299 content increases in observations and models. This highlights the need for reliable cloud  
300 property observations — particularly cloud phase — in sufficient spatial and temporal  
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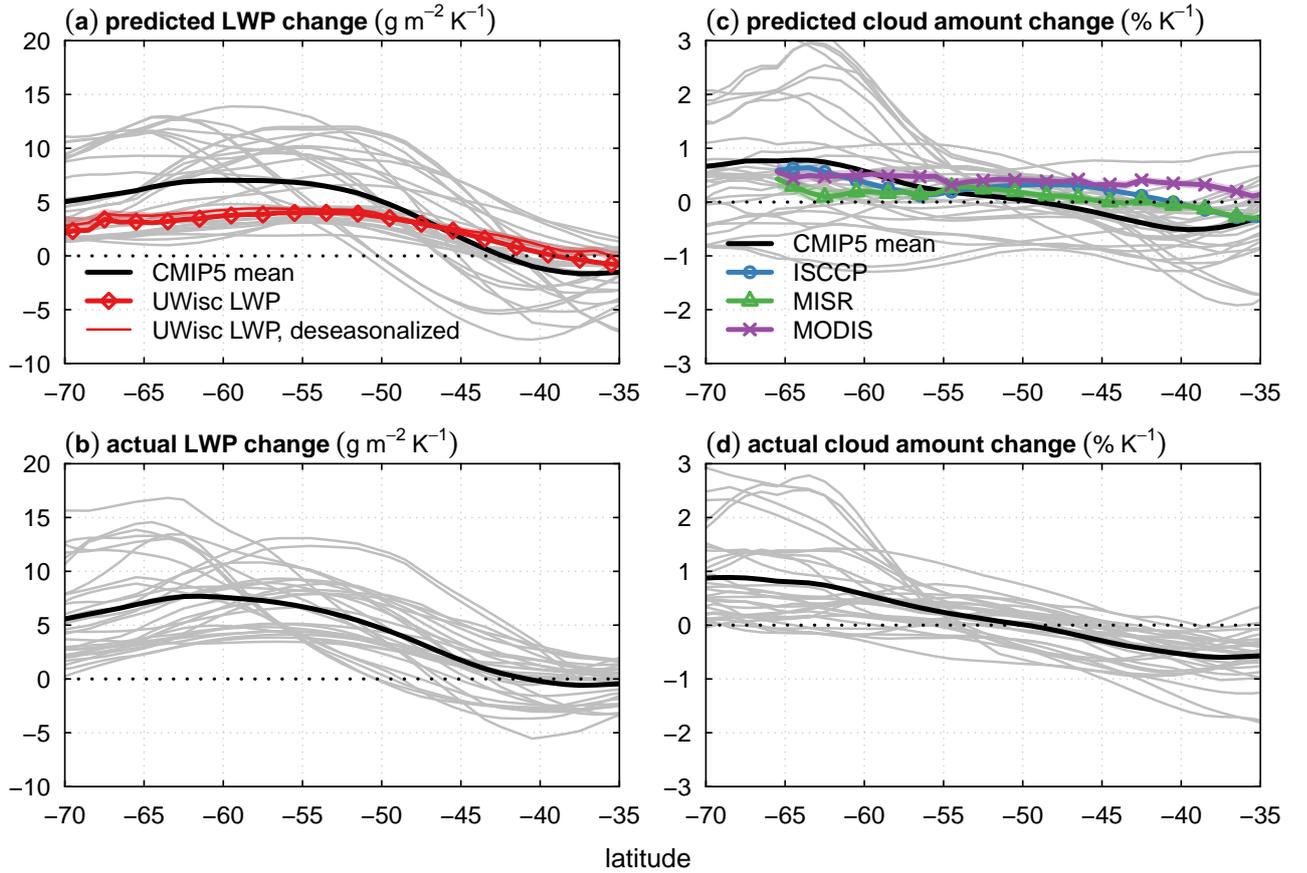
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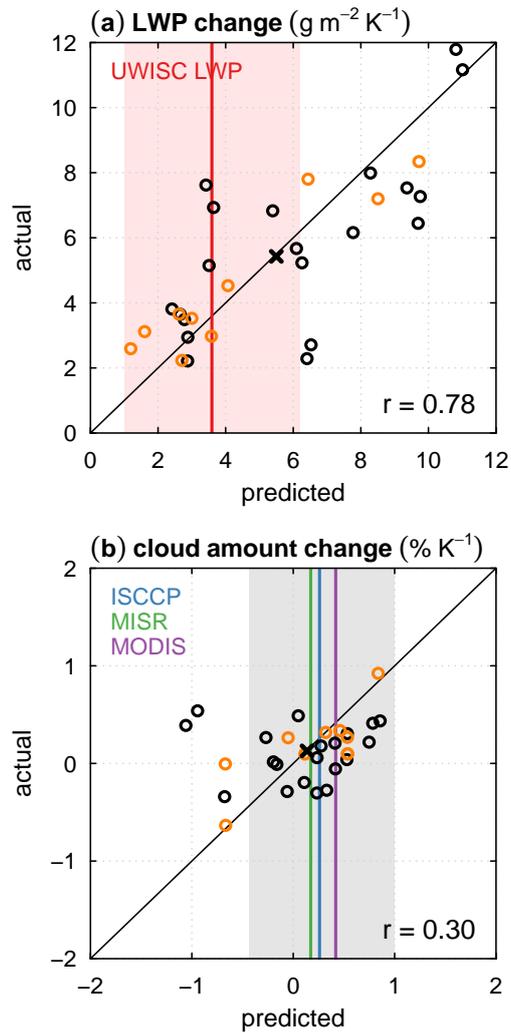
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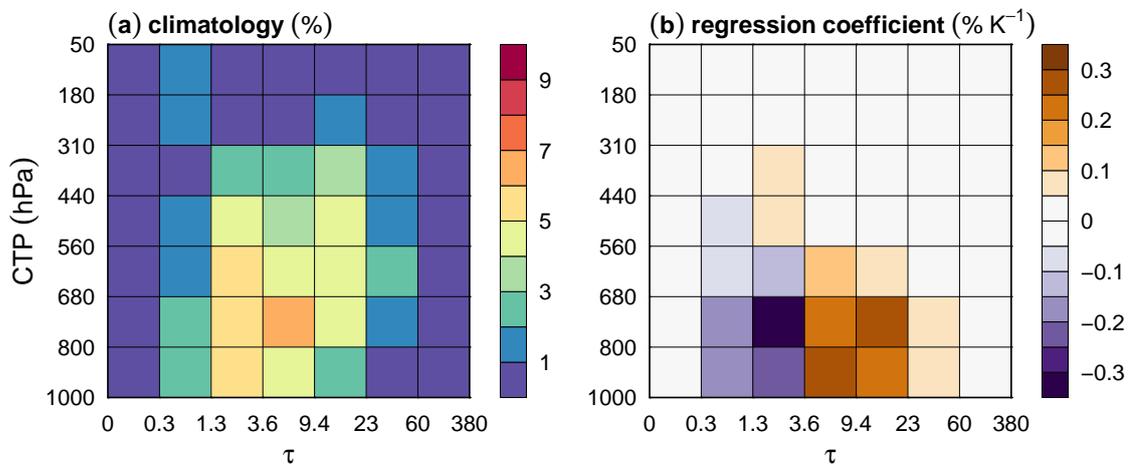
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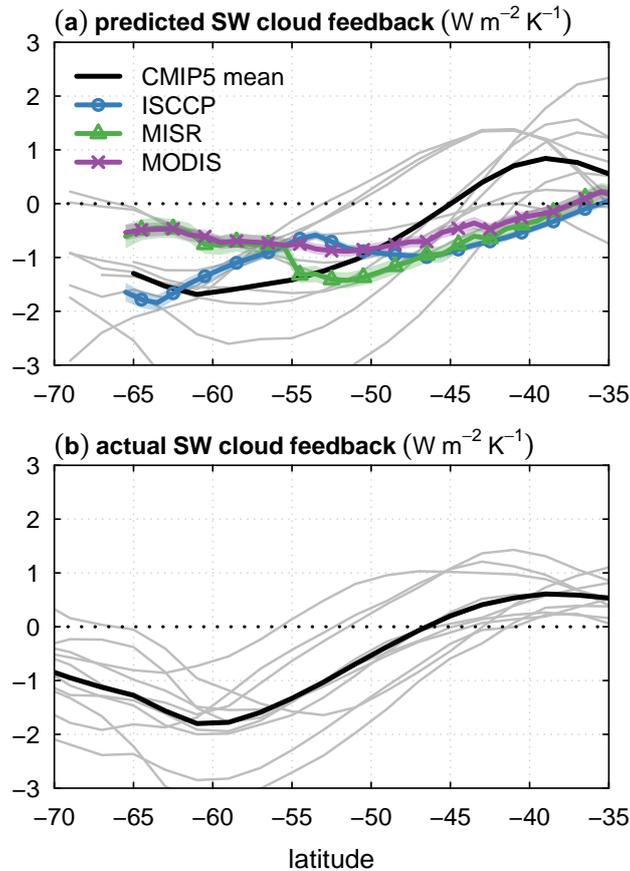
**Figure 1.** Predicted and actual change in liquid water path (LWP) and cloud amount in CMIP5 models and in observations. The predicted changes (a, c) are based on regressions onto low-level (500–850 hPa) temperature (see Methods); the actual changes (b, d) are calculated as 2081–2100 minus 1981–2000 in the RCP8.5 experiment, normalized by the low-level temperature change in each model. Observed cloud amount is obtained by integrating the cloud property histograms over all cloud top pressure and optical depth bins. Grey curves denote individual models, the thick black curve represents the multi-model mean, and colored curves correspond to observational data sets. For observations, pale color shading denotes the 95% confidence intervals based on a two-sided significance test for the regression slope; these confidence intervals are often not visible due to their narrowness.



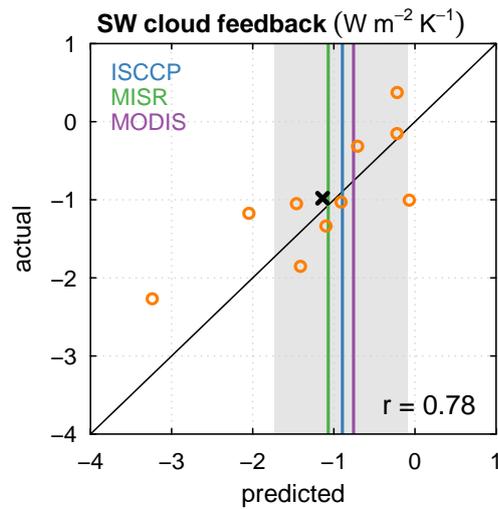
**Figure 2.** Scatterplots of actual versus predicted (a) LWP change and (b) cloud amount change, both averaged over  $45^{\circ}$ – $60^{\circ}$  S. The black crosses mark the multi-model mean. Vertical colored bars denote the predictions associated with observational data sets, and orange open circles indicate the models with cloud property histogram data that were used for the SW cloud feedback calculation in Figs. 4 and 5. Shading represents the 95% uncertainty interval for the mean observational estimates, calculated as described in Text S3.



**Figure 3.** ISCCP cloud fraction histograms binned by cloud top pressure (CTP) and optical depth ( $\tau$ ). The mean climatology is shown in (a) and the regression coefficient on low-level temperature in (b), both averaged over the 45°–60° S latitude range.



**Figure 4.** Predicted and actual SW cloud feedback in models and observations. Curves are defined as in Fig. 1. The predicted cloud feedback is obtained by multiplying cloud-radiative kernels [Zelinka *et al.*, 2012b] with cloud fraction histograms regressed on temperature. The actual cloud feedback in (b) is calculated using APRP with RCP8.5 data (see text), and includes cloud adjustments to  $CO_2$  forcing. The feedbacks are normalized by the local low-level temperature change rather than global-mean surface temperature. For observations, pale color shading denotes the 95% confidence intervals based on a two-sided significance test for the regression slope. In panel (a), missing data at high latitudes result from the exclusion of ice- and land-covered grid points from the regression analysis.



**Figure 5.** Scatterplot of actual versus predicted SW cloud feedback, averaged over  $45^{\circ}$ – $60^{\circ}$  S. Vertical colored bars denote the predictions associated with satellite observations. Grey shading represents the 95% uncertainty interval for the mean of the three observational estimates (Text S3). The black cross marks the multi-model mean. The orange open circles are as in Fig. 2.

**1 Supporting Information for “Observational evidence  
2 for a negative shortwave cloud feedback in mid to  
3 high latitudes”**

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#### 4 **Contents of this file**

5 1. Text S1 to S3

6 2. Figures S1 to S4

7 3. Table S1

#### 8 **Introduction**

9 In addition to four supplementary figures and one table listing the models used in our  
10 analysis, in this Supporting Information we provide information on the data sets used in  
11 this study (Text S1), discuss some possible sources of error in the satellite retrievals that  
12 may affect our results (Text S2), and describe the methodology to calculate observational  
13 confidence intervals (Text S3).

#### 14 **Text S1**

15 In this section we provide additional details on the observational products and model  
16 data sets used in this study.

17 The ISCCP, MISR, and MODIS data sets are optimized for comparison with GCM  
18 satellite simulators, and are provided in a format identical to standard CMIP5 model out-  
19 put by the Cloud Feedback Model Intercomparison Program (CFMIP; [http://climserv.  
20 ipsl.polytechnique.fr/cfmip-obs/](http://climserv.ipsl.polytechnique.fr/cfmip-obs/)). The ISCCP data are a simulator-oriented ver-  
21 sion derived from the D1 data set, covering the period July 1983–June 2008 (25 years).  
22 For MODIS, Collection 5.1 data are used, covering March 2000–February 2015 (14 years),  
23 and including both Terra and Aqua satellite retrievals. Finally, MISR Level 3 data are  
24 used, covering June 2000–May 2013 (12 years). Since MISR retrieves cloud top physical  
25 height rather than pressure, we interpolate the MISR cloud histograms from height to

26 pressure coordinates, while conserving total cloud amount. In doing this we assume a  
27 simple relationship between pressure  $p$  and height  $z$  given by  $p(z) = p_0 \exp(-z/H)$ , with  
28 surface pressure  $p_0 = 1000$  hPa and scale height  $H = 8$  km.

29 In CMIP5, the most commonly available satellite simulator is the ISCCP simulator  
30 (variable clisccp), for which 13 models have provided output in either the historical or  
31 AMIP experiments. As in *Zelinka et al.* [2012], we verify correct simulator implementation  
32 by comparing the simulator-based total cloud amount (obtained by adding over all cloud  
33 top pressure–optical depth bins) with the actual total cloud amount output by each model,  
34 and reject models for which the two measures of global-mean climatological cloud amount  
35 differ by more than 10 percent. Based on this test, we use cloud histograms from 10  
36 models (crosses in Table S1); the rejected models are IPSL-CM5A-LR, IPSL-CM5A-MR,  
37 and MIROC-ESM.

## 38 **Text S2**

39 In this section we discuss possible sources of error in the satellite retrievals, their possible  
40 impact on our results, and our data analysis choices to mitigate such impacts.

41 ISCCP retrievals suffer from substantial inhomogeneities in time (due to changes in  
42 instrumentation) and space (due to the different spatial coverage of geostationary and  
43 polar-orbiting satellites) [*Evan et al.*, 2007], and particularly in high latitudes unphysical  
44 spatial inhomogeneities in cloud properties are obvious even upon simple visual inspection  
45 of the raw data; such spatial artifacts are possibly related to polar-orbiter swaths [*Norris*,  
46 2000]. For this reason, we doubt that the large increase in the negative feedback poleward  
47 of  $55^\circ$  S reported by ISCCP is physical (blue curves in Figs. 4a and 6a).

48 Furthermore, optical depth observations are affected by errors due to temporal varia-  
49 tions in viewing and illumination geometry, whose effects depend on instrument technol-  
50 ogy and retrieval algorithm. Of particular relevance to instruments using solar reflectance  
51 such as ISCCP, MODIS, and MISR are the potentially large optical depth biases at the  
52 highest solar zenith angles (i.e. at high latitudes during winter), as discussed in section  
53 3 [Loeb and Davies, 1996]. For MODIS, these solar zenith angle biases are large and  
54 well-documented [Seethala and Horváth, 2010; Grosvenor and Wood, 2014; Lebsock and  
55 Su, 2014].

56 We have attempted to mitigate the impact of such systematic errors by omitting re-  
57 trievals at solar zenith angles larger than  $60^\circ$  for MODIS and MISR (see Methods). If such  
58 values are not excluded from the analysis, the SW cloud feedback predicted by MODIS  
59 becomes weaker (less negative) by  $0.25 \text{ W m}^{-2} \text{ K}^{-1}$  between  $45^\circ$  and  $60^\circ$  S, but in MISR  
60 it strengthens by about  $0.67 \text{ W m}^{-2} \text{ K}^{-1}$  (not shown), leading to larger disagreement  
61 between observational data sets. Since optical depth is overestimated at the highest solar  
62 zenith angles (and therefore at the lowest temperatures, opposing the positive optical  
63 depth–temperature relationship), one would expect the illumination angle effect to lead  
64 to an underestimation of the negative optical depth feedback, as for the MODIS results.  
65 This not being the case for MISR suggests other effects may play a role; for example,  
66 complex interactions between viewing and solar zenith angle are known to affect MISR  
67 retrievals [Liang and Di Girolamo, 2013].

68 Although we excluded retrievals at the highest illumination angles, it is likely that  
69 seasonal variations in solar zenith angle still affect the optical depth–temperature rela-

70 tionships in our results. Repeating the cloud histogram regressions with deseasonalized  
71 data — and thus completely filtering out the possible signal due to solar angle varia-  
72 tions — we still obtain increases in cloud optical depth and a negative SW feedback,  
73 but the results are very noisy and the historical relationships are poor predictors of the  
74 forced response in models (not shown). However, since the observed LWP–temperature  
75 relationships remain nearly identical when the seasonal cycle is removed (Fig. 1a), we  
76 are confident that the cloud optical depth increase with warming is not an artifact of  
77 seasonally-dependent satellite retrieval biases.

78 While the lower LWP sensitivity to temperature in observations compared to the multi-  
79 model mean would suggest that the optical depth feedback is much too negative in models  
80 (cf. Fig. 1a), the LWP observations are not exempt from errors either. *Lebsock and Su*  
81 [2014] have shown that on average, microwave LWP retrievals (similar to the data set  
82 used in this study) tend to substantially overestimate the true LWP, with errors mainly  
83 stemming from ambiguities in the detection of cloudy pixels and in the partitioning be-  
84 tween cloud and precipitating water. In light of these biases, *Lebsock and Su* [2014]  
85 concluded that observational LWP climatologies remain too uncertain to effectively con-  
86 strain climate model behavior. Thus, we caution that the observed LWP–temperature  
87 relationship shown here should not be interpreted in a strictly quantitative sense.

### 88 **Text S3**

89 Observational confidence intervals are derived from the uncertainty in the prediction of  
90 the models' responses based on historical regressions on local temperature. The uncer-  
91 tainty in the prediction of cloud property responses is thus taken as proportional to the

92 standard deviation of the residuals relative to the one-to-one line in Figs. 2 and 5. As-  
93 suming that the residuals are normally distributed, the half-width of the 95% confidence  
94 interval is then simply equal to 1.96 times the standard deviation of the residuals. The  
95 confidence intervals are calculated relative to the mean observational estimate in Figs. 2b  
96 and 5.

97 Our calculation is similar in concept to the estimation of prediction intervals in linear  
98 regression [see e.g. *Wilks*, 2006, Eq. 6.22], but we do not include the terms related to the  
99 uncertainty in the intercept and slope of the relationship, since the expected relationship  
100 between predicted and actual responses is known (the one-to-one line). Furthermore, here  
101 the residuals are calculated as orthogonal departures from the one-to-one line, rather than  
102 as vertical departures as in least-squares regression.

103 The confidence intervals calculated in this manner do not include the uncertainty in the  
104 estimation of the observational regression slopes, which is very small in our analysis owing  
105 to the robustness of the relationships. More importantly, our calculation also excludes  
106 the uncertainty associated with errors in the satellite retrievals, reflected in the differences  
107 between the various observational estimates. We believe this uncertainty is mitigated, but  
108 not fully suppressed, by taking the mean of the three observational estimates as the most  
109 likely observational value.

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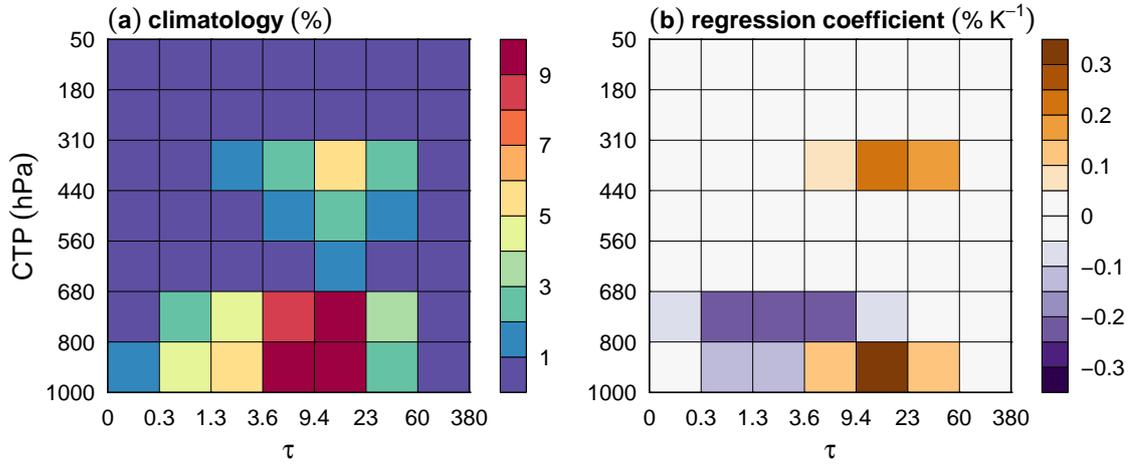
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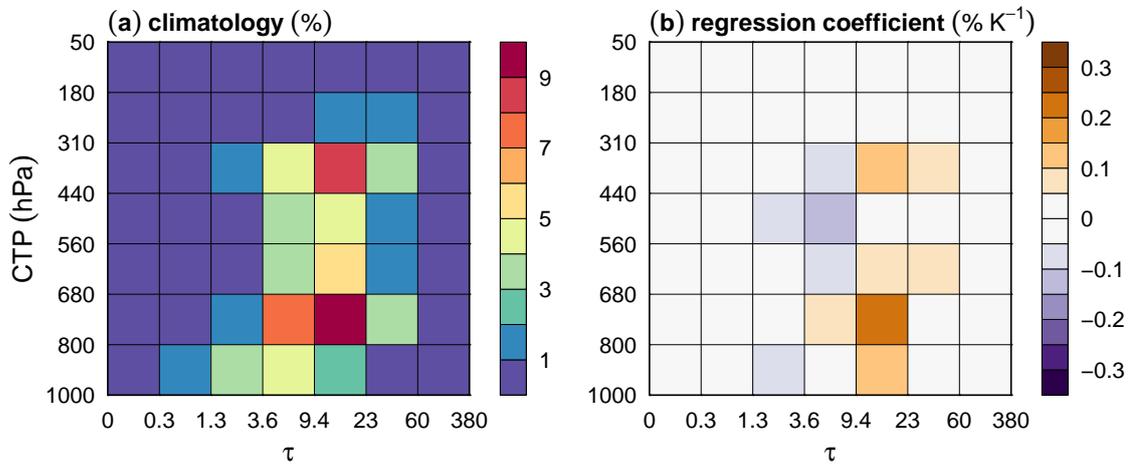
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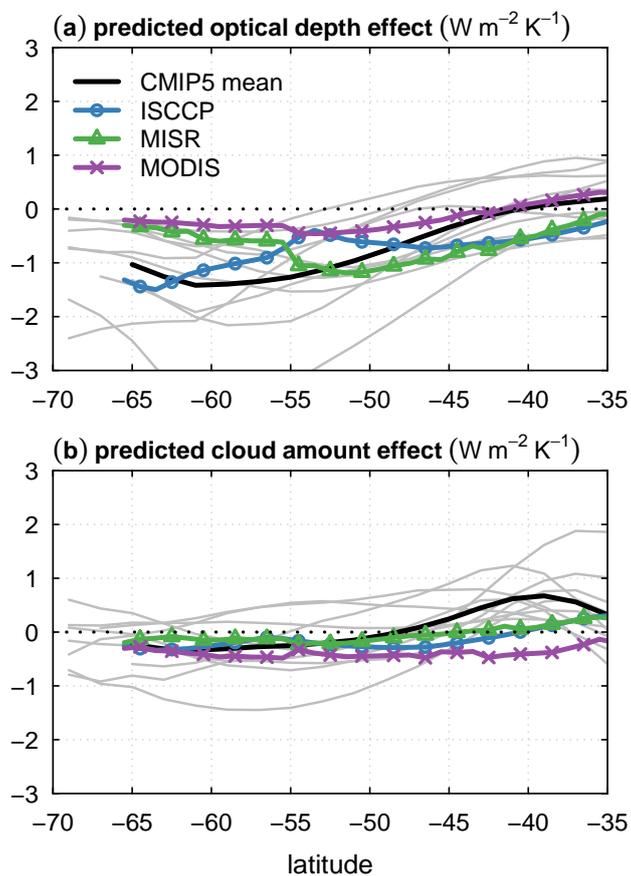
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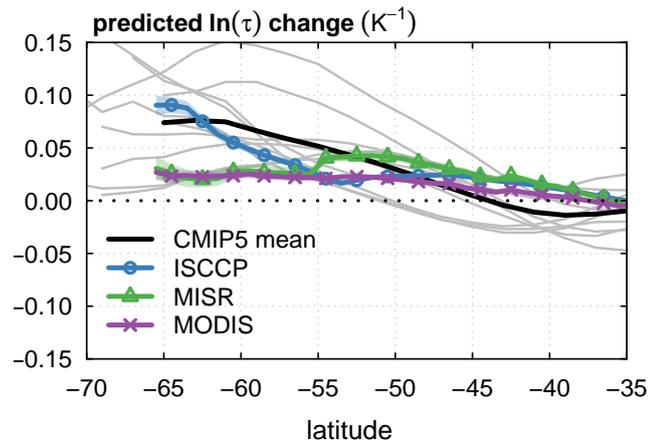
**Figure S1.** As in Fig. 3, but for MISR. The cloud amount values have been remapped from cloud top height bins to cloud top pressure bins (see Text S1) for consistency with the cloud-radiative kernels.



**Figure S2.** As in Fig. 3, but for MODIS.



**Figure S3.** As in Fig. 4, but partitioning the cloud feedback into effects of (a) optical depth changes and (b) cloud amount changes, following *Zelinka et al.* [2013]. Confidence intervals are omitted.



**Figure S4.** Predicted change in the natural logarithm of optical depth,  $\ln(\tau)$ , calculated by regressing  $\ln(\tau)$  onto low-level temperature. Here optical depth is simply calculated as the weighted mean of the 7 optical depth bins in the cloud property histograms, using the total cloud amount values in each bin as weights.

**Table S1.** List of CMIP5 models used in our analyses. A cross ( $\times$ ) indicates the models that provided historical or AMIP cloud property histograms using an ISCCP simulator. All 30 models listed below provided LWP, total cloud amount, as well as temperature data for the historical and RCP8.5 experiments. For each model, only the first ensemble member is used.

	Model name	historical	AMIP
1	ACCESS1.0		
2	ACCESS1.3		
3	BCC-CSM1.1		
4	BCC-CSM1.1(m)		$\times$
5	CanESM2	$\times$	
6	CCSM4		$\times$
7	CESM1-BGC		
8	CESM1-CAM5		$\times$
9	CNRM-CM5		$\times$
10	CSIRO-Mk3.6.0		
11	FIO-ESM		
12	GFDL-CM3		$\times$
13	GFDL-ESM2G		
14	GFDL-ESM2M		
15	GISS-E2-H		
16	GISS-E2-R		
17	HadGEM2-CC		
18	HadGEM2-ES	$\times$	
19	INMCM4		
20	IPSL-CM5A-LR		
21	IPSL-CM5B-LR		
22	IPSL-CM5A-MR		
23	MIROC5	$\times$	
24	MIROC-ESM		
25	MIROC-ESM-CHEM		
26	MPI-ESM-LR	$\times$	
27	MPI-ESM-MR		
28	MRI-CGCM3	$\times$	
29	NorESM1-M		
30	NorESM1-ME		